



THE ASTRONOMER'S YEARBOOK 2019

Complete observing guide to the year

Detailed star charts for January to December

The seasons' essential constellations up close

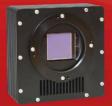
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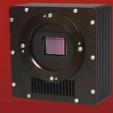


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Gaze up on a clear night and wonder awaits. Planets, stars, nebulae, constellations, galaxies, meteor showers... all waiting to be discovered. Whether it's the sparkling spectacles of crisp autumn and winter evenings, or the glorious

displays of a spring or summer night sky, we've got Inside this Yearbook you'll find guides and star

charts for every month of the year - when and where to find the finest sights in the ever-changing view above our heads.

It's not just a matter of finding them, of course. The real reward is in understanding and appreciating what you're seeing – and this guide will help you do that. Top astronomy experts share their tips, not just on spotting well-known superstars of the skies, but on lesser-known but enchanting targets too. We've also set some challenges to help you hone your stargazing skills - give them a go and see how you get on.

All you need to spot many of the marvels on offer are your eves, but things get really exciting when you add in some magnification, even if it's just a good pair of binoculars, Or, if you're in the market for some top-end kit, you'll find reviews and advice on equipment here too. There's also some superb astrophotography that's sure to whet your appetite, whether you're just starting out with a DSLR or itching to see what your CCD setup can achieve.

While there's plenty to keep you busy at night, we've got your daytimes covered too with our step-by-step DIY guides to making your own astronomy accessories.

With this Yearbook you'll know what to look forward to and what not to miss in 2019. Enjoy!

Chris Bramley, Editor













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The winter constellation Orion forms one of the most famous and easily recognisable patterns in the night sky

The familiar winter constellation of Orion holds many surprises for imagers who want to delve a little deeper, says Will Gater

There are few constellations that grab the attention quite like that icon of the winter heavens, Orion. The alittering bright stars. the instantly recognisable 'belt' and the many glowing nebulae scattered within the Hunter's boundaries all make Orion a wonder to behold on a frosty, dark night. But the constellation is also a rich hunting ground for astrophotographers seeking captivating targets of many kinds. In this article we're going to explore some of the

different ways Orion's splendours can : be captured on camera, from a simple nightscape that conveys the naked-eye view to advanced CCD imaging techniques that can reveal the constellation's extraordinary deep-sky features. And hopefully, by the end of this piece, you'll agree with us that no matter how many times you catch sight of the Hunter, you'll always find something new to inspire you and test your astrophotography skills. >



ABOUT THE WRITER Will Gater is an astronomy journalist author and presenter. Follow him on Twitter at @willgater or visit willgater.com

EXPERIENCE LEVEL Beginner to intermediate WHAT YOU'LL NEED A DSLR

bridge camera and a stur raphic tripod. A wide kit lens for the kind that comes with most DSLRs) will be perfectly sufficient. More experienced astrophotographers may also want to use a portable tracking mount to capture longer exposures.

There's something tremendously evocative about glimpsing the bright stars of Orion over a wintery landscape - or towards the end of the autumn months, just as the nights start to get longer and colder - so in this first part we're going to look at how to shoot a 'nightscape' that attempts to capture some of that magic.



STEP 1 Make a conceptual plan Thinking about the emotions you want to

convey or elicit with your shot can help you to plan a powerful picture, and it'll inform every stage of the photographic process. For example, if you wanted to evoke the harsh iciness of winter observing you might shoot Orion over an isolated, leafless tree in a barren landscape, and process the image in such a way as to create a hard contrast between land and sky.



STEP 4 Compose with the landscape and sky conditions To compose your

niahtscape vou can take short, very high ISO test exposures to show you the balance and positioning of foreground and sky, and any structures or

landscape features in frame. Try to use the foreground – trees, buildings, etc – to lead the viewer's eve toward Orion. Sometimes clouds can be used as a framing device too, and thin cloud can even 'bloat' and enhance the colours of bright stars.



STEP 2 Select your focal length or a prime lens

Once you've thought about what atmosphere you want to capture with your image, you can select the focal length you'll be shooting at. A typical kit lens set to around 24mm, or an equivalent prime lens, provides a wide field of view for Orion on a camera's sensor, allowing you to fit in the brighter central stars and the Hunter's fainter outlying 'arms'.



STEP 5 Set the exposure length, aperture and ISO

When shooting, keep the lens aperture wide open (lowest f-stop), though some lenses will perform better when reduced a few stops. Experiment with the ISO and exposure length until you're happy with the look You may need to use an exposure that very slightly trails the stars in order to define the foreground.



STEP 3 Focus the view

Next, focus the view. Some cameras have a live preview function that can be zoomed onto a suitable star, giving you instant feedback as you make slight focusing adjustments. With Orion there's no shortage of bright stars that can be used for this. Repeat the process a few times - checking that the star is as small as possible - so you're certain the image is as sharp as it can be.



STEP 6 Process your image

When processing nightscapes, reducing the noise in the image and bringing out foreground detail are the main challenges. As long as you shoot in RAW format, modern image-processing software is well-equipped to handle these tasks. In Photoshop or GIMP you can correct the colour balance, and use the 'Curves' tool to bring out star fields and improve overall contrast and definition. >

Orion overview

Where to find some of the constellation's best imaging targets

Bright orange star

A striking reflection nebula with billowing clouds of dark, dusty nebulosity

The Horsehead and Flame nebulae A spectacular pairing of iconic nebulae with contrasting pinkish-red and gold colours

M42 and the

Running Man nebulae
The magnificent Orion
Nebula (M42) and its
beautiful blueish companion

Bright blue-white star

The Witch Head N A ghostly reflection nebula near Rigel

Barnard's Loop
A huge arc of glowing red gas visible in deep wide-field images

Far and WIDE

Reveal the hidden delights lurking within Orion with the help of long-exposure, wide-field imaging

EXPERIENCE LEVEL Intermediate WHAT YOU'LL NEED A DSIR, a racking mount and either a relatively long focal length comera lens (between 100 and 300 mm focal length on a full-format DSIR) or a short focal length refractor. You could use a CCD camera, but the field of view produced by your setup will need to be at least 5" across or you'll need to mosaic.

One of the things that makes Orion so attractive for astrophotography is the diversity of deep-sky objects within its borders, from pinkish-red star-forming regions to blue-tinted reflection nebulae.

The proximity of these targets to one another means that long-exposure widefield imaging of Orion can produce some spectacular compositions. Not only do such wide-field images show the positions of objects such as the Orion and Horsehead nebulae in relation to one another, but they can also reveal the rarely seen fainter surroundings of objects that are usually given the 'close-up' treatment, such as the aforementioned nebulae.

A DSLR with a long focal length lens and mounted on some form of equatorial tracking mount is probably the simplest setup with which to get started in wide-field imaging. Unlike most deep-sky imaging, wide-field deep-sky astrophotography generally doesn't require autogulâng, as it's possible to capture good data with unguided subexposures of jixs a minute or two.

With fast prime lenses and those relatively short exposure lengths, you may be surprised at how easily you can pick up some of Orion's most recognisable deep-sky objects. For the best results, capture multiple sub-exposures (as well as dark formers and flat fields) and then calibrate and stack them using software such as the free DeepSkyStacker, before final enhancements in your preferred image-processing software.



Colourful captures

With the right setup, you can show that Orion is more than just white stars against a black background

EXPERIENCE LEVEL Beginner
WHAT YOU'LL NEED A basic DSIR or
bridge comers fitted with a lens that
allows manual focusing (some compact
digital comeras will also work, depending
on the lens/focusing mechanism they
use). You'll also need a photographic
tripod and your camera will need to be

The colour variation of Orion's bright stars is one of the most captivating things about the constellation, yet it can be tricky to capture these wonderful hues as the chromatic obseration in some camera lenses overwhelms the true star colour. One method for showing the tints of stars such as Betelgeuse, Rigel and W Orionis is to manually defocus the image. It's a technique that was made famous by the renowned astrophotographer David Malin many years ago. You can use this method with a wide lens for a fast long lens) and static tripod, as long as you use short exposures — a second or so in the case of a longer lens. All you do

is frame the star (or constellation), defocus the lens a little by hand and capture an exposure, usually at a mid-ta-high level ISO setting, In the two composite images below we focused on Betelgeuse and Rigel. We captured a number of exposures and in between each one we defocused the lens a bit more. Then we combined them into one frame using processing software. It's a very artificial composition, but it does give a flowar of one of the things that makes observing and immain 20 in special.





3 4

▲ Combining progressively defocused images of Betelgeuse (left) and Rigel (right) will highlight the contrasting colours of Orion's brightest stars

Portrait of a stellar nursery

Capture the Orion Nebula's ethereal pink swirls of gas and dust that are giving birth to new stars

EXPERIENCE LEVEL Intermediate to advanced

WHAT YOU'LL NEED A small refractor or Newtonian telescope carried on a motorised tracking mount, plus a monochrome CCD camera (and a computer to control it) with a set of LRGB imaging filters and a filter wheel. For exposures of more than a few minute it's also a good idea to use an autoguiding system alongside the above, though this is not absolutely necessary.

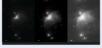
There are few greater tests of a deep-sky astrophotographer's skills than the magnificent Orion Nebula, M42. Among the many challenges it provides are the faint outer regions of the nebula that can be lost in processing, or simply not picked up at all during the imaging process, and its dazzlingly bright core that requires careful planning to capture. In the step-by-step guide below we've described the basic process of how to go about shooting M42 with the kind of setup you might typically have if you're starting out in CCD imaging - that is, a monochrome CCD camera and a set of LRGB filters (luminance, red, green and blue) with which to make a full-colour image.





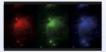
STEP 1 Set up and polar alian accurately

Once you've got your equipment set up. spend some time finessing the polar alignment of your mount. This is so you'll be able to get the longest unquided exposures your mount is capable of before the stars drift out of position - this is especially important if you're not using autoguiding equipment.



STEP 2 Capture different length **luminance** exposures

Use short, 'binned', test exposures to compose the image. Then take three groups of exposures through a clear luminance filter; short ones for M42's bright core, longer ones for the main body and, for the faint outer regions, as long as your unguided mount can manage without the stars 'wandering' (usually several minutes)



STEP 3 Get the RGB colour

When you've got 10-15 sub-exposures for each of the three groups of luminance data. you can move on to capturing the colour data through red, green and blue filters. Capture at least 10-15 images per colour channel - aim for an exposure length similar to your shots of the main body of M42 with the luminance filter.



STEP 4 Take dark frames and flat fields

After capturing each 'LRGB' channel, carefully stretch a clean white pillowcase or t-shirt over the scope aperture (without touching the lens) and illuminate it with a torch before taking an image. This is a flat field, which records image artefacts such as vignetting and dust on the optics. Also take a set of dark frames if the data from your CCD needs them.



STEP 5 Stack and calibrate the data

You should now have six sets of sub-exposures: three luminance groups of varying exposure lengths and one for each of the RGB channels. Load them into your preferred astronomical imageprocessing software (for example, DeepSkyStacker) and use the flat fields and dark frames to calibrate them before stacking those calibrated sets into six images



STEP 6 Combine the three luminance images

Bring the three luminance images into layers-based image processing software, such as Photoshop or GIMP. With each image in a separate layer, erase the overexposed portion of the long-exposure image so that the 'main-body' exposure shows through - do the same for the main body layer so the core shows clearly. Merge the layers.



STEP 7 Add the colour and make final

processing adjustments

Next, place your red, green and blue filtered images in their respective colour 'channel' in a new image file. Copy the resulting full-colour image, as a separate layer, into the luminance file created in Step 6 and turn its blending mode to 'Colour'. Lastly, make any final image tweaks to your taste

For those of us unable to see the aurorae, there's another way to experience the effects of particles streaming from the Sun on Earth's atmosphere

Measuring

Creating a homemade magnetometer allowed Stuart Green to keep a record of space weather, despite cloud and his southerly location



ABOUT THE WRITER A composite materials engineer, Dr Stuart Green is a keen solar astronomer and nthusiast

Earth is bathed in a constant stream of energetic particles originating at the Sun. This solar wind ebbs and flows, and occasionally explodes in a coronal mass ejection, throwing billions of tonnes of our star's plasma into space, which impacts on Earth's magnetosphere - our protective magnetic bubble. The result is bright and beautiful aurorae, the luminous splendour of which is our only visual confirmation of this Earth-Sun connection. For those fortunate enough to witness this spectacle, the lasting impression is one of awe at its magnificence. For the rest of us, such events are lost save for the images available online, through which we can only experience the aurorae vicariously

A few years ago I was looking for a way to establish a degree of connectivity with the Sun that was otherwise unavailable to me in my lower latitude location in the UK. That's when I thought about building a magnetometer. Not only does the solar wind create the aurora, but in the process of being deflected by our protective magnetosphere, it also imprints its signature upor our protective bubble. This is detectable at any point on our planet with a magnetometer, which picks up the magnetic signature as it fluctuates

orientation of the passing plasma.

Magnetometers can be purchased commercially, of course, but my interest lay in building one for the fun of it and for the challenge. The basic scheme of the design is a highly sensitive magnetic sensor, an ultrasonic emitter, an ultrasonic-to-audio frequency converter and a computer with a sound card and Spectrum Lab audio spectrum analysis software for data logging.

A stable environment

At the heart of this setup is the magnetic sensor. Called a fluxgate, these sensors can be extremely sensitive and are perfectly suited to measuring the tiny perturbations in the local magnetic field I used provides an output frequency that varies according to the strength of the magnetic field.

As well as responding to the magnetic field, the output frequency also changes with temperature. so the sensor has to be located in a temperaturestable environment. My sensor is buried in my garden about 0.5m below the surface. It is also garden about o's moen's mis annue. It is sup-protected by a waterproof housing made from 40mm diameter plumbing pipe fitted with end caps sealed using suitable solvent.

The sensor is positioned almost perfectly level on a poured concrete base and points along

stable to avoid spurious readings, so mine is linked to a linear voltage regulator that converts 9V .



► DC from a mains transformer to a stable 5V DC supply.

When pointing east west, the output from the fluxgate sensor is a train of +5V square wave pulses at a frequency in the range of 60-70kHz, which is significantly above the operating frequency of any standard computer sound card. So I turned to equipment intended for zoology and connected the output of the magnetic sensor to a bat detector with an ultrasonic transducer emitter.

Bat detectors are designed to convert ultrasonic bat calls into audio frequencies that we can hear. Commonly this is accomplished by a process called heterodyning, in which an internally generated, tuneable reference frequency is mixed with the

varying ultrasonic input signal from the bat to create audible sound at a frequency equal to the difference between input and reference frequencies.

Frequency range

For the purposes of magnetometry, this frequency difference should be tuned to a single audible tone that can be recorded with a standard computer sound card. I found 3-4kHz to be a suitable range. The other advantage of using a bat detector is that the audio output is of good fidelity, meaning that subsequent data analysis can be accomplished at high resolution. For this project, I used a detector with an internal crystal oscillator for precision and to minimise any frequency drift. This detector is

▲ The magnetometer is made up of a magnetic







Temperature stability is vital, hence the magnetometer (left) is housed in a vacuum flask (centre) which is kept inside a cool box (right)



The sound of oace weat

Once the data from the solar wind hitting Earth's magnetosphere has been collected, it's possible to create an audio file from it

Taking things a step further with an online algorithm, it's possible to convert the output magnetic field strength data into sounds using the MIDI (Musical Instrument Digital Interface) protocol. In this process, each field strength value is converted to a MIDI value representing a particular musical note, which is then fed into music production software to prepare an audio track.

MIDI allows any digitised instrument to be assigned to play the note, although my preference is to use a synthesised wind sound to represent the solar wind. In reality, the data relate directly to local magnetism, not to the solar wind, yet to hear the sounds generated by such a method is evocative as one imagines the stream of particles pushing against our protective magnetic field, occasionally entering to

Once the magnetic declination data generated by my magnetometer was sonified, it could be synchronised with a timelapse video of the aurora borealis to create an audio-visual experience similar to this one: bit.ly/2n9ohFo. The correlation between the video, created by Kai-Marius Pedersen in Tromsø, Norway, and the audio is not perfect - probably because of the physical separation between it and the sound recording in the UK – but there is a satisfying degree nevertheless. More sonified geomagnetic data can be found at bit.ly/2n9rHc1.

"I now have a connection with the cosmos, with data streaming into my PC constantly, capturing the ebb and flow of the solar wind"

> usually powered by a set of batteries, but as that only provided a battery life of less than 24hours. the batteries were swapped for another stabilised DC feed from the 9V mains transformer.

Gatherina data

With the detector set up I was ready to log data. and for this I used free software called Spectrum Lab, written by Wolfgang Buescher (www.gsl. net/dl4vhf/spectra1.html). This enables the data stream to be displayed as a continuous chart as the day progresses. The software can also send measured frequency data to a computer hard drive at preselected intervals, together with a time and date stamp, building up a database throughout an entire geomagnetic event. Later, this database can be exported as a CSV file and copied into an Excel spreadsheet for further analysis. Any time period between logged events can be selected; I use a

one-second cadence for high resolution work and a 150-second cadence for standard resolution work.

Analysing the data involves converting frequency to magnetic field strength (more accurately magnetic flux density in nano-Tesla) using conversion factors provided by the sensor manufacturer, and then charting the result as a function of time.

As it turned out, my magnetometer needed a few refinements to make it more consistent with professional data, in particular better temperature control of the bat detector. The output from the detector was being significantly impacted by small ambient temperature fluctuations that were imprinting on the output frequency. It was only when the detector was placed in a temperaturedefined environment (inside a vacuum flask, inside a cool box fitted with a vivarium heater mat at a controlled temperature) that the output met the professional readings.

The output from such a relatively simple device is remarkably consistent with data generated by professional aeomagnetic monitoring stations such as Eskdalemuir, operated by the British Geological Survey, and its sensitivity and resolution certainly compete favourably for the hobbyist with the output from magnetometers costing considerably more.

Having established my homemade magnetometer I now have that connection with the cosmos that I was looking for, with data streaming into my PC constantly, capturing the ebb and flow of the solar wind and occasional coronal mass ejection that buffets our planet.



Take it from me

Experienced astronomers reveal to **Will Gater** the secrets they've learnt from their hundreds of hours of stargazing and astrophotography

When you're looking to brush up your stargazing skills, whether it's observing, imaging or processing techniques, or deciding which equipment is worth buying, one of the best ways is to pick the brains of knowledgeable friends who can offer you advice, or even show you how something's done. That's why we've spoke how

four experts to find out the most useful things they've learnt over the years, what mistakes to avoid and the most important things they think every beginner should know. Over the next few pages you can find the distillation of those conversations and we hope it'll be just the thing to quide you through those crucial early steps.



DEEP-SKY ASTROPHOTOGRAPHER

Sara started out in astrophotography in 2010. With the help of astronomy forums and books, she learnt imaging techniques and is now one of Europe's leading deep-sky astrophotographers. From her observatory in Spain she regularly produces stunning pictures of galaxies, nebulae and star clusters



tips for deep-sky imaging

"Think about what you're imaging and it there's anything else nearby that you can include that will add to the image."

ilters, pick your arget appropriately secause some targe example – aren't going to respond at all to narrowband imaging."

"Create a setup that's as simple as possible – something with a

"I would say the biggest challenge in deep-sky imaging is getting used to the mount, so that you're able to polar align. Without polar alignment you're not going to get longer exposures - certainly if you're not autoguiding

"People seem to think there's a rite of passage in which you go from a DSLR to a one-shot colour CCD camera and then to a monochrome CCD. I don't necessarily think people should feel that they have to progress like that. If you've got the inclination, and you think it's going to be a hobby that you'll enjoy, then buy secondhand equipment and just jump in at the deep end.

"Your neighbour's probably not going to be doing astrophotography as well so they can't stand there and help you, but that doesn't mean you're on your own. Internet forums are fantastic - I've gained so much information from them and made so many friends, it's unbelievable.

"I've learnt that refractors are far easier to use and, for me, are the better option, I started off with refractors before moving on to a Schmidt-Cassegrain telescope, a Ritchev-Chrétien and then an optimised Dall-Kirkham. But I've gone back to a refractor because it's just easier. I don't want to





waste time faffing around; I just want to switch everything on and go.

"As you get more experienced, the way you compose a shot changes. You learn to balance things out, look ground and see what else is there. I think that's really quite important, because that ultimately gives you more interesting pictures.

"When it comes to executing that composition, I've found plate solving software is incredible. It'll take a picture of the sky, compare it to various catalogues that it's got on record and then tell you exactly where you are. For example, say you want to photograph the Pinwheel Galaxy, M101, and have framed it in the top left-hand corner. You can tell the software this is where you want it to go and it'll take a picture, then say, 'Actually you're 6,000 pixels away, and move the frame by 6,000 pixels. It'll take another picture, then say, 'You're still three pixels out' and make the three-pixel adjustment. After a few shots you'll be in exactly the right place night after night, and can even go back to it year after year. With an image from last year, I can plate solve it then ask the software to point the scope at the same spot and that's where it goes." >

[►] Sara's image of M13, Cluster in Hercules



Richard

VISUAL OBSERVER

It was the Moon landings that steered Richard towards astronomy. One of his early optical instruments was a pair of Second World War binoculars, which he recalls using to observe the Andromeda Galaxy. Today he's an experienced visual observer, astrophotographer and an expert in meteor astronomy



"Beginners often ask me what telescope to buy and I reply: 'Don't buy a telescope, Just get a pair of binoculars and learn how to use them.' If what you see through the binoculars is impressive, then consider investing in a telescope. Astronomy is impressive, but it's not necessarily a visual spectacle - it's understanding what you're looking at that's the key. Binoculars are enough to do that, "If you do decide that you want a telescope, don't buy one until you've tried viewing through the telescope you intend to buy and have experienced the sort of views you're going to see with it. Judging purely on the pictures of the views a telescope can offer often leaves people disappointed by what they actually end up seeing. "With telescopic observing I think you need to be interested in something. If you're just looking

at Messier objects you'll run out of targets quite quickly. I got interested in variable star observing because a member of my local astronomical society encouraged me to view things over a long term. With variable stars you don't see what's going on during a single night. You have to build up a series of observations to understand what's happening. So I think you've got to do a bit more than simply spotting objects, otherwise you're not really getting the benefit. It's nice to get the Messier objects, but what do you do once you've seen them all? This is where things like variable stars and comets come in

"One of my big regrets is that I don't write things down enough. I don't keep a diary and I think, looking back, there are a lot of things I've forgotten. The background ground an observation is often just as interesting - not so much what I saw, but the adventure of observing it. Those are just memories for me now and eventually I'll lose them."





loved observing



PROTECT YOUR EYES

Solar observing can be dangerous. Never view or photograph the Sun without using a certified solar filter that has been sourced from a reputable supplier and which is used and installed as per the manufacturer's safety instructions. Also, be sure to remove or cap any finderscopes your telescope may have prior to observing

SOLAR OBSERVER

Lyn has been fascinated by astronomy since she was seven, but it was a programme on solar observing presented by Sir Patrick Moore that sparked her specific interest in our nearest star. Today she's the director of the British Astronomical Association's Solar Section



Lyn's favourite solar phenomena

oler regions c Sun's 'surface e Jun's 'surtace' at appear as dark lemishes on the right solar disc."

sma leaping of scopes and r systems."

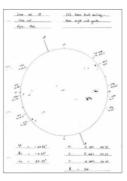
t appears as a y bright patch pically near an tive region or large hspot). You can w these through ecialist hydrogen-ha-filtered solar

► Sketching sunspots gives

"When you start observing it can feel like there's a lot of pressure on you to become an expert on everything straight away, and trying to do that isn't what you want to do at first. You need to enjoy it. You need to start simple, learn a little bit and then progress onto the next stage.

"Never look through a telescope at the Sun unless you've got a solar filter on it. You can look through a telescope as long as you have a good-auglity solar filter that uses Baader AstroSolar Safety Film - that gives you a nice. white, silvery image. Alternatively you can buy a Type 2 glass solar filter and put that over the objective end, which will give you a kind of yellowy-orange Sun.

"The good news about solar astronomy is that small is beautiful. Normally in astronomy you're trying to capture as much light as possible, so you need a big aperture. With solar astronomy it's the





opposite: you've got too much light so a nice smallaperture telescope is what you're after.

"I sketch every single observation, which I think really helps. If you're trying to count sunspots or anything, you need to draw it. It helps your brain sort out what's there and helps you count. It's all too easy to just look at the full picture, but when you're drawing something you have to concentrate more on the detail.

"Bring the sunspot, filament or whatever you're looking at into the centre of the evepiece. Then you'll see it far better. Wait for that moment of clarity, of fine seeing. There's that fraction of a second where it's so clear you see everything and then it's gone again. That's why observations take a while: you need to keep waiting for that moment of perfection and then quickly sketch it down.

"The most difficult thing for me to comprehend as a beginner was the orientation of the solar disc. It can be so simple to work out east from west when you know how, but I wish I'd had someone to tell me when I first started. Just let the solar image drift through the eyepiece and the first limb to slide out of view is the western limb. That's a really good thing to get in your head." >



WRFNCE

HIGH-RESOLUTION LUNAR AND PLANETARY IMAGER

Today. Pete is an experienced imager of the Moon and planets, but his first shot was taken by pointing a digital camera down the eyepiece of a telescope. He now uses high-frame-rate cameras and a large-aperture Schmidt-Cassegrain telescope to capture spectacular pictures of our celestial neighbour



looks on your imaging chip. The longer the focal length of the telescope, the larger the object you're imaging will appear. You can't change the physical focal length of your telescope but you can change the effective focal length by using an optical amplifier such as a Barlow lens or a Powermate. For example, a 2x Barlow will give you a focal length that's effectively two times your telescope's natural focal length.

"But beware: there is a danger associated with doing this. If you increase the image scale too far, the planet or the Moon will just look like mush because you need perfect conditions to do that. You're not using the camera's pixels at their optimum efficiency.

"One of the best ways of processing an image - whether it's deep-sky, planetary or lunar - is to fiddle about with it to create your final image, then walk away from it for 20 minutes or more. When you come back and sit down in front of it, your image will either look good or it'll look awful, Don't worry if it looks awful, everyone produces bad images from time to time. When you're processing. it's all too easy to tweak an image to death, but if it's not working, just walk away. When you come back you'll see the processed result for what it is."





for lunar and planetary imaging 1. "Make sure your

2. "Make sure the scope has cooled emperature with the utside before

3. "Make sure your focus is spot on. Go through focus, out t e focus posit oks like "

"A big danger in astronomy, as in many walks of life, is that once one method of imaging the Moon and planets becomes second nature, you tend to assume that it's the only method there is and that you shouldn't vary from it. That's wrong. You shouldn't assume that the way you're doing something is the only correct way - you should always experiment and try different things. And don't be afraid of making mistakes.

"I made plenty of mistakes when I started out. My first images of Mars looked like somebody had taken a photograph of a baked bean, because it showed very little detail - it was just an orange blob. It needed that push to increase the image scale in order to bring out some of the surface details. Image scale is basically how big something



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Stellarium, created by French programmer Fabien Chéreau in 2001, is a ubiquitous astronomy tool for the digital age





Astronomer Will Gater shows you how to master the free planetarium software Stellarium and make the most of some of its more advanced features

There are few pieces of software that have made such a contribution to amateur astronomy as the free planetarium program Stellarium. Available from stellarium.org for Windows, Mac OS X and Linux systems, it is popular with beginners and advanced stargazers alike. Indeed, not only can it do everything you'd want from a planetarium program – that is, simulate the night sky and the positions of planets and other celestial bodies at a given date and time - but it's also packed with other features that are tremendously useful for experienced observers and astrophotographers. Here we take a look at the basics of the

program - so that if you're new to it you'll have an introduction to how to use the software - before then exploring further by examining Stellarium's plua-ins and other useful tools that even we. with decades of stargazing experience under our belt, come back to use time and time again. We suggest you fire up the program while you read this so you can experiment as we go. >



The basics of Stellarium

The symbol-based menus may seem arcane at first, but this guide will quickly get you up to speed



The main screen

1 There are two main toolbars. The bar along the bottom shows key information such as the location, the field of view (FOV) and the time. Placing your cursor over this bar will raise it to reveal buttons that can togale display settings, such as constellation lines and star labels. This bar is also where some plua-in buttons will appear.

2 On the bottom-left is another toolbar to access further configuration windows, to select the location you're observing from. set the date and time, and tailor what's shown on screen. The Search tool and Help window are also here.

3 If you have the Oculars plua-in (see page 26) you will see its icons on the top-right.

4 If you left-click on a celestial object from the main screen, the object will acquire a small rotating target around it (stars) or a bouncing blue frame (other objects). Information about the object will appear at the top of the screen, including its name, magnitude and catalogue numbers. Press F2 to customise this display in the 'Information' tab.

Setting the date and time

When Stellarium first opens, it synchronises with your computer's date and time. If you open it during the day it'll show a simulated daytime sky: to see the stars come out you'll need to advance the time. But the program handily lets you specify a time (say 11pm) that Stellarium will default to every time you open it: press F2 to access this on the 'Navigation' tab. In general use, if you want to display the sky for a specific date and time, simply click the clock on the left toolbar and input the desired values.





Expert tips Stellarium allows

a). We like to set the 'Absolute and the Relative otorealistic look

Setting the 'viewing' location

Selecting the location from which you want to view the sky in Stellarium is simple. Click on the compass icon in the left-hand toolbar or by pressing F6 to bring up the 'Location' window. In that window, you can then choose a location either by clicking on a

world map then selecting from a list of nearby places or by typing in a town or city, or inputting an exact longitude and latitude. You can even select an extraterrestrial viewing location if you want to simulate the sky from other bodies within the Solar System.

Exploring Stellarium's sky

To explore the sky in Stellarium you can drag it round with your mouse or use the arrow keys on a keyboard. To zoom in and out use 'ctrl' (or 'cmd' on a Mac) plus the up and down arrows, or you can use a roller wheel on a mouse. To zoom in on a specific target first left-click on it, then press the spacebar to centre it before using the zoom keys.



How to tweak what Stellarium shows

Pressing F4 brings up a 'Sky and viewing options' window. Here you can change the size of the stars, adjust the number of labels and limit star magnitudes (useful for decluttering the chart) and simulate light pollution. From the 'DSO' tab you can choose which catalogues of deep-sky objects are flagged, while the 'Markings' tab allows you to switch between various celestial markers lines and coordinate grids. The 'Landscape' and 'Starlore' tabs let you change the foreground scenery and alter which culture's star names and constellation patterns are used. .

Using the search function

Stellarium's search menu can be accessed by pressing F3. Pressing the 'return' key on a successful search will immediately centre that object on the screen.

Other stargazing software choices

There are many other sky charts programs available if Stellarium doesn't quite fit your virtual stargazing requirements:

Sky Safari: skysafariastronomy.com Starry Night: www.starrynighteducation.

Cartes du Ciel Skychart: www.ap-i.net/ skychart/en/start

Google Sky: www.google.com/sky Redshift: www.redshift-live.com/ext/en ESA Sky: sky.esa.int

Sky Map: www.sky-map.org



Advanced features

Here we delve deeper into Stellarium to look at some of the useful plug-ins for imagers and observers



Simulating eyepiece and camera fields of view

Stellarium comes with an extremely useful feature - the 'Oculars' plug-in - that allows you to simulate the field of view of a given eveniece and telescope combination, as well as overlay the field of view of a camera and optics setup onto the sky. Of all the plua-ins in the program this is the one we use the most, and its ease of use makes it a superb tool for planning astrophotos. The plua-in can be activated by accessing the 'Configuration' window using the F2 button, then clicking on the 'Plug-ins' tab. Scroll down the menu on the left to 'Oculars', then click on the 'Load at startup' button, When you reopen Stellarium, you'll see a new toolbar has appeared in the topright of the screen.

symbol) brings up Ocular's configuration window where you can list what telescopes, evepieces and cameras you own. The second button activates a camera field of view tool that places a red box onto the simulated sky, representing the boundary of the view of the specified equipment. The third icon switches on a circular 'target' pattern that's used in Telrad finders, Swapping between various equipment combinations in the program is

The rightmost of these (a spanner

easy: click the double arrows next to each piece of equipment listed in the box under the toolbar, which appears when either the eveniece field of view or camera field of

view tools are in use. The camera field of view can even be rotated by increments so you can experiment with the ideal framing of a potential astro image.



Expert tips

tints the whole Stellarium window red, to help preserve your dark adaptation if you're

Using and updating the 'Solar System Editor'

Stellarium doesn't just simulate the positions of stars, deep-sky objects, the planets and the Moon. It also shows where numerous asteroids and comets are located at any given moment. New comets and asteroids are discovered frequently and you can keep the program's databases up to date using the 'Solar System Editor'. Press F2 to find it in the plug-ins menu.

Clicking the 'Configure' button on the Solar System Editor page brings up a new window with a 'Solar System' tab The button to 'Import orbital elements in MPC format' lets you import orbit data for asteroids and comets from lists maintained by the Minor Planet Center: simply select a list and click 'Get orbital elements'. Choose which bodies you want Stellarium to plot before clicking 'Add objects'. You can then search for these objects using Search (F3).





Showing satellites

Among Stellarium's many excellent advanced features, the 'Satellites' plug-in is likely to be the one of most interest to watchers of humanity's orbital outposts. Once enabled via the 'Plua-ins' menu, it plots the positions and movements of a huge number of satellites. This is not just interesting to look at, it also creates a superb planning tool for photographers

who enjoy taking pictures of things like the International Space Station, geostationary satellites and iridium flares - especially when used with the camera field of view function of the 'Oculars' plug-in, For example, if the ISS is going to make a pass over the location that Stellarium is simulating, the program will show it as a bright point of light moving against the

stars, just as you would see it in real life. Stellarium can even simulate the Station going into - or emerging out of - Earth's shadow, and the behaviour of iridium flares, with the satellite momentarily brightening dramatically before fading. To make the path of the ISS even clearer, we advise checking the 'Orbit Lines' box on the plugin's configuration page.





Celestron's Wi-Fi-controlled Astro Fi 5 comprises a 5-inch (127mm) Schmidt-Cassegrain telescope on a single fork arm attached to an aluminium tripod. The telescope has a focal length of 1,250mm giving a focal ratio of f/10, which means this is considered a 'slow' system, best suited to observing planets, lunar and bright deep-sky targets. It looks very smart in its Astro Fi livery, while the optical surfaces come with Celestron XLT

multi-coatings for a crisper quality view. The single fork arm has a Vixen-style mounting saddle to which you attach the telescope using a chunky restraining bolt; it was pleasingly easy to install on the tripod. A padded power pack, which takes eight AA batteries (not included), is provided and you can also use an optional power tank. The power switch has a small red LED, which is handy. A StarPointer red dot finder, star diagonal and two eyepieces complete the system. The 25mm eyepiece gives a magnification of x50 while the 10mm eyepiece provides x125, a good range for the kind of targets you'd use this system for. The Wi-Fi/smartphone control of

the system worked well. Powering up the Astro Fi 5 and connecting via the SkyPortal app (iOS and Android), we were soon aligned using the on-screen instructions. We could then use the app to explore our targets or manually slew the scope using the onscreen arrows.

The Astro Fi 5 is a decent, easy-to-use system aimed at beginners who want to enjoy the night sky using their smart devices. It certainly managed to fulfil its criteria and can be recommended.

- Price £599
- Optics 127mm SCT compound • Focal length 1,250mm, f/10
- . Mount Wi-Fi Go-To altaz single-arm mount
- . Ports Power connector, two aux ports. integrated Wi-Fi adapter
- Tracking rates Sidereal, lunar, solar trackina
- . Tripod Adjustable tripod with
 - accessories tray including rubberised smartphone holder
- Extras StarPointer red dot finder. 25mm and 10mm 1.25-inch fit
 - eyepieces and star diagonal
- Weight 7kg
- . Supplier David Hinds Ltd
- Tel 01525 852696 · www.celestron.uk.com

Objective lens caps

Poorly fitting objective caps are about as useful as a chocolate fireguard. Luckily, the plug-in objective caps on these binoculars are excellent; they provide a tight-fitting seal and can only be removed by using the tabs provided. If you lose them, it won't be because they fell off in a field somewhere.

The soft rubber eye cups easily conform to the shape of your eye sockets, making them comfortable to use. They fold down to give an effective eye relief of 15mm, which we found to be adequate to allow us to wear spectacles for observing



The case has side rings so you can use it like a shoulder bag, but also has three rings on the back, so you can use it like a rucksack and distribute the weight evenly across your back rather than all on one shoulder.



optical tubes, the central mounting bar has a sliding mounting position which enables you to easily balance the binoculars on the mount. This design also keeps the binoculars' centre of mass close to the mounting head's altitude fulcrum; this reduces balance asymmetry at different altitudes.



Can these budget-priced big binos compete with more expensive models?

Bia binoculars can be great astronomical instruments, but they are often prohibitively expensive. So our interest was naturally piqued when some budgetpriced, 80mm aperture binoculars came onto the market

We tested the Opticron Oregon Observation 20x80s under a variety of sky conditions, mounted on a photographic tripod with a fluid video head. Stars snapped to a good focus, which looked consistent over the central 80 per cent of the field of view. There was some offaxis chromatic aberration on the Moon's terminator and limb, but generally both colour correction and colour rendition were good. It was possible to easily distinguish the colours of Alderamin (Alpha (a) Cephei), Zeta (t) Cephei, Delta (δ) Cephei and Mu (ω) Cephei.

These binoculars come into their own on extended deep-sky objects. The

Orion Nebula was bright, showing good structure, and we could distinguish all four stars of the Trapezium (Theta (0) Orionis). The galaxies M81 and M82 were easy to see and the differences between them were obvious. The Andromeda Galaxy. M31, nearly filled the field of view and we could detect the more abrupt cut-off owing to the dust lane on the near side. The normally difficult-to-see M33 and M101 galaxies were easily spotted when they were high in the sky, especially when we used averted vision. Although the Ring Nebula, M57, did not appear as a distinct ring, it was easy to detect and had the appearance of a disc with a darker middle.

getting damaged or scuffed. It also helps to reduce glare

from objects just outside the field of view.

If you fancy trying a larger-thanstandard pair of binoculars without breaking the bank, the Opticron Oregon Observation 20x80s should certainly be on your shortlist, particularly if you're new to binocular astronomy. These binoculars are pleasant to use, provide crisp, clear views, have no alarina faults and also come with a five-year UK guarantee to provide significant peace of mind.

- Price £149
- · Optics Fully multi-coated
- Aperture 80mm
- Magnification 20x
- Prisms BAK 4
- . Angular field of view 3.26°
- Focusing Zeisscentre focus • Eve relief 17mm
- Interpupillary distance 56-77mm
- · Weight 2.2kg
- Supplier Opticron
- Tel 01582 726522 • www.opticron.co.uk





Altair Starwave 70 EDQ-R quad apo imagina refractor

A four-lensed, lightweight scope designed for easy, wide-field astro imaging

The first thing we noticed about the 70 EDQ-R is how light it is: with a Canon EOS 70D DSLR camera attached it weighed just 2.1kg. This really does make it a graband-go telescope for even the smallest of mounts and ideal for travel.

Build quality in general is very good, and it comes with quite a lot as standard. The tube is made from lightweight alloy and it comes with a good set of adjustable tube rings on a dovetail bar. A quality rack and pinion focuser with nice, smooth movement adds to the quality feel. The only thing that let it down was a loose lens cap that kept falling off.

The first target we imaged was the Andromeda Galaxy using a finder-guider configuration on a portable tracking mount set on a tripod. The camera was a Canon DSLR setup using the basic

EOS capture on the computer. We took 20 three-minute images which captured lots of detail in the dust lanes with round stars and vibrant colours. For such a short amount of imaging time we were happy with the end result. The same process was repeated on a subsequent night with the Pleiades as our target, again revealing plenty of structural detail.

Chanaina to CMOS cameras was simple and with M42-threaded extension rings on the camera it was easy to focus on the screen. When you're using a mono camera, there's ample space to attach a small filter wheel in the imaging train if you need it. Images of the Moon with a Hypercam 183C showed a lot of detail for the aperture.

We concluded that the 70 EDQ-R is a very good addition to anyone's imaging setup. With the right camera attached it can fit some very large objects into its field of view, and makes it easy to capture good images with minimal effort.

- Price £1.150
- · Optics Air-spaced quadruplet Aperture 70mm
- Focal length 350mm, f/5
- . Focuser CNC rack and pinion with fine focus
- Extras Tube rings, dovetail bar.
- finder bracket
- · Weight 1.4kg . Supplier Altair Astro
- Tel 01263 731505
- www.altairastro.com

Lightweight aluminium tube rings and a Vixen dovetail bar are included. The dovetail bar has an elongated slot at one end for easy adjustment of the rings. This makes balancing the position of the telescope easy on any motorised mount. The bar has a thread to mount on a tripod and ball mounts.

Dew shield

A sturdy built in, fully retractable dew shield helps keep dew down to a minimum in average temperatures. It is coated black on the inside to stop any stray light entering the telescope. The dew shield has adjustment screws to make a firm fit to the tube.

Finder shoe

Having a finder shoe is a nice touch as they're often not included on smaller refractors. Based around a Sky-Watcher finderscope, the shoe is quite generous in size and doesn't interfere with the rotation of the camera. It can also be used to mount a small guidescope or camera.



A high-quality rack and pinion focuser with a micro focuser for fine

adjustment makes focusing easy. It can be fully rotated allowing the telescope to be mounted on a range of equipment without the focus knobs getting in the way. The focuser can support even the heaviest of cameras with nice smooth adjustment.

The M42-threaded rear adaptor is fully rotatable and makes framing targets really easy. All locking pins are nylon coated for secure locking. On the inside of the flange are adjusters to allow for sensor tilt. With the tilt adjuster set to zero the back focus is 66.78mm.

Equatorial mounts require polar alignment to avoid star trails and field rotation during long exposures. The MiniTrack is supplied with a very simple polar finder - an 8mm sighting tube that fits into a clip on the end of the drive unit, which works well enough.

There is no declination axis on this type of mount so the camera and lens have to be attached to a ball-andsocket head so that they can point to anywhere in the sky. The beautifully engineered Omegon ball-and-socket head makes pointing at celestial objects very straightforward.

The ball-and-socket head has a well-designed saddle clamp with a single, very positivefeeling hand bolt. This saddle clamps a minidovetail bar forming an excellent quick-release mechanism for easy attachment and removal

A seven-level spring tensioner is built into the mounting head to help alleviate any off-balance pressure being applied to the clockwork drive mechanism.



1/4-20, UNC-threaded mounting hole right in the centre of the casting. This allows the MiniTrack to be attached to a wide range of tripod heads and equatorial wedges.



Omegon MiniTrack

LX2 trackina mount

The power to track stars is in your own hands... literally, with this clockwork device

Historically, the drives and gears used to motorise telescope mounts have been known as 'clock drives' because the original mechanisms had much in common with early clocks. The Omegon MiniTrack LX2 brings this older technology bana up to date by using a clockwork motor to power a compact equatorial tracking mount. This elegant solution makes for an ultra-portable mount with no power requirements other than a gentle winding every hour or so

Assembly is very quick indeed. You start by attaching the MiniTrack to your own tripod's pan-and-tilt head or precision altaz head so that you can tilt the tracker to polar alian it. You then attach the ball-and-socket head to the mounting platform. Finally, you attach the camera and lens to the quick release dovetail bar, which we easily managed even in the dark. To start the imaging session, you turn the large disc at the foot of the mount

a maximum of one rotation to wind up the clockwork motor - it's as simple as that. Once the motor's running, you can loosen the ball-and-socket head to allow the camera lens to point to the celestial object you want to photograph and, after focussing, image capturing can begin.

We used an external intervalometer attached to our Canon 450D and 28mm wide-angle lens set to capture a continuous set of three-minute exposures with the camera in 'hulh' mode. We used our own hot shoe-mounted red dot finder to aim the lens in the general direction of the bright star Sadr (Gamma (v) Cvani) in Cyanus, and captured images until the motor ran out, which was exactly an hour later. The resulting star shapes were impressive, showing no sign of trailing even though we were using the maximum exposure length recommended for our kit.

We would recommend the MiniTrack to users of any experience level. More than

just a aimmick, this clockwork marvel is a simple and effective means of mounting basic equipment for wide-field imaging.

- Price £139
- · Payload capacity 2kg
- · Latitude adjustment Adjustable wedge
 - or pan-and-tilt tripod head required
- Tracking rates Sidereal
- · Power requirements Clockwork, wound by hand
- · Tripod Not supplied
- · Extras Ball-and-socket head, polar finder tube. ¼-20 thread adaptor.
- · Weight 774g with ballhead
- · Supplier Astroshop.de
- Tel 00 49 8191 940491
- www.astroshop.eu

Sky-Watcher StarGate-500P SynScan Dobsonian

Bigger appears to be better when it comes to telescopes



'Freedom-Find'.

The StarGate-500P is a half-metre class instrument and weighs 90kg - make no mistake, it's not for the faint hearted.

Using it to take a tour of some of the best deep-sky targets the sky has to offer, we were wowed by the views. In twilight, the double star Albireo was stunning in both evenieces. Later under dark skies, the M35 star cluster in Gemini filled the view of the 28mm and we noted coloured stars scattered throughout it. We picked up the normally fainter nearby cluster NGC 2158, which was bright and a mass of stars.

Objects that appeared as a smudge of light in smaller instruments took on new life with the 500P, such as M1, the Crab Nebula, which showed as a strongly mottled oval with undulating edges and clear views of the central stars.

Turning to galaxies, we couldn't help but gasp when we homed in on the Cigar Galaxy, M82, with the 10mm eveniece and it displayed lots of the dust features we normally associate with photographs. Dropping down to the 28mm eyepiece, we could just fit its neighbour M81 in the view with a clear spiral nature to it. We homed in on Uranus and were rewarded with a lovely pale green ball instead of the usual view of a small disc. We even picked out its five moons, although Miranda was still a challenge.

Overall, the StarGate 500P is an impressive system; societies and clubs that have sufficient funds and want a real deep-sky light bucket with Go-To functionality will find that this is a great instrument to invest in. Just remember to keep a step ladder with it, as looking up towards the zenith puts the eveniece more than 2m above the around!

- Optics 508mm (20-inch) primary mirror
- Focal length 2.000mm f/4
- Mount Dobsonian Go-To Altazimuth · Ports Power connector, SynScan AZ
- hand controller
- Database 42,900+ objects
- · Tracking rates Sidereal, lunar, solar tracking and Freedom-Find Dual-
- Encoder system • Extras 10mm (1.25-inch), 28mm (2-inch) evenieces, 9x50 straight-
- through finderscope, fabric shroud · Weight 90kg
- Supplier Optical Vision Ltd www.opticalvision.co.uk

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Month by month guide to

What to see in the night sky throughout the year by Pete Lawrence

36 January

Optimal conditions for the Quadrantids and a total lunar eclipse

40 February

The year's largest supermoon and our best sight of planet Mercury

44 March

A chance to enjoy galaxy-laden Leo as the spring equinox arrives

Fingers crossed for the zodiacal lights on Moonless evenings

54 May

Dark skies could make for exciting Aquariids and M5 is well placed

58 June

Hercules and Ophiucus dominate June's skies

64 July

The nights may be short but Saturn's rings shine bright in July

68 August

All eyes on Lyra, Cygnus, Aguila and bright Jupiter

72 September

Plenty of evening viewing and Neptune skips past Phi Aquarii

78 October

Stunning views as several galaxies hecome favourable

82 November

A transit of Mercury and a lunar occultation are two rarities to relish

86 December

Taurus delights as Orion provides ample hunting ground

Key to monthly star charts Use this key to get the most out of each month's detailed star chart



The Small 47 Tucanae ▷

Michael Sidonio, Canberra, Australia. 23 October 2017

HOTSHOTS WINNER JANUARY 2018

January

Optimal conditions for the Quadrantids and a total lunar eclipse are among the highs

1 JANUARY

Comet 46P/Wirtanen – a naked eve possibility

2 JANUARY

Venus, Jupiter, Mercury and a 6% waning crescent Moon visible one hour before sunrise

3/4 JANUARY

The peak of the Quadrantid meteor shower with no Moon to mar the view

Total lunar eclipse from 02:36 UT until 07:48 UT

22 JANUARY

Venus and Jupiter appear 2.5° apart in the morning sky



A Venus and Jupiter will appear close to one another in the morning sky

Janua

The New Year is heralded with a superb view of the night sky, made all the better at the start of the month by the new Moon occurring on 6 January. With clear weather, the bright, vibrant stars of winter will offer an irresistible invitation to explore the cosmos. The most obvious pattern is Orion the Hunter, seven bright stars forming a very familiar pattern. Orion is rich in deep-sky wonders, the most obvious starting point being the Hunter's Sword, a line of faint stars hanging vertically below his very obvious Belt

The sword comprises stars and deep-sky objects. The most obvious is M42, the Orion Nebula. This is visible using binoculars or a telescope and appears as a swirling mass of glowing gas with a small cluster of stars at its heart called the Trapezium because of its four brightest members. A dark lane separates M42 from comma-

shaped M43 to the north. Both ends of the Sword are marked by open clusters, NGC 1980 to the south and NGC 1981 to the north. Turn to page 48 to find out about more of Orion's wonders. Another somewhat overlooked object is

M41, an open cluster

is visible in the

MOON PHASES

Key stages in the monthly cycle









VISIBLE PLANETS

Where to spot the planets this month



month 30 minutes before sunrise

Reaches areatest western (morning) elongation on 6 January. Close to Jupiter on 22 and 23 January.

Through a telescope, evening object Mars appears small with a distinct gibbous phase.

JUPITER Bright morning planet in Inhiuchus Close to Venus on

SATURN May be glimpsed in the morning sky with Venus and Jupiter at end of January.

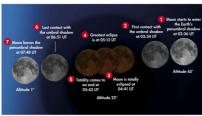
LIRANUS Well placed in the evening sky best seen at the start of January.

Neptune requires binoculars. Enirly well positioned it will be best at the start of January.

in Canis Major, the Great Dog. Starting at Orion's Belt, extend the line made by the three stars down to the left as seen from the UK. They point at the brightest nighttime star of them all, Sirius (Alpha (α) Canis Majoris). M41 is located 4" below Sirius. Placing Sirius at the top of the field of view of average binoculars. M41 should be somewhere close to the bottom. The cluster contains around 100 stars and occupies an area equivalent to the apparent size of the full Moon.

Star strings and a grinning cat Another interesting open cluster lies to the east of

Orion. Starting from Bellatrix (Gamma (v) Orionis) extend a line through Betelaeuse (Alpha (a) Orionis) for about 1.3 times the distance, bendina slightly south as you go. Here you should be able to see NGC 2244, a small open cluster in Monoceros that bears a similarity

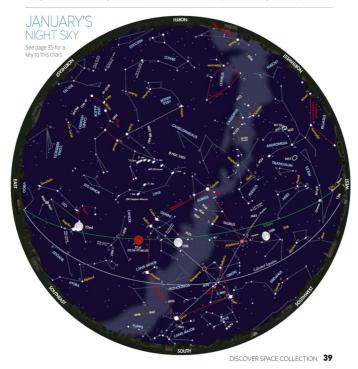


▼ As long as the

to the number six on a die. Although this looks like a regular open cluster, it's special in that it marks the heart of the beautful Rosette Nebula, a large area of circular nebulosity. The Rosette is normally the domain of long exposure astrophotography but can be seen visually through a telescope fitted with an Oll filter.

As this month is turning out to be something of a cluster-fest, it's also worth mentioning M35 at the foot of the Twin Castor in Gemini. This is a lovely object to view through a small telescope and contains many 'star-strings', patterns of stars which appear to form curving lines. Further north, following the line of the winter Milky Way through into Auriga, a famous Messier trio of clusters rewards a high-altitude binocular sweep. The three are numbered out of sequence running south to north as M37, M36 and M38, with M38 being part of the amusing asterism known as the Cheshire Cat.

January will be an exciting astronomical month with plenty going an Highlights include forounable circumstances for the annual Quadrantid meteor shower peak on the night of 3/4 January (see page 62 for radiant position), a total celligase of the Moon in the early hours of 21 January and a close pairing of Venus and Jupiter on the morning of 22 January.







Didier Rediger-Lizloy/ David Attié/Stefan Shramm/e-Eye remote telescope hosting service. Spain, 25-30 October 2017

HOTSHOTS WINNER FEBRUARY 2018

February

The shortest month sees the year's largest 'supermoon' and our best sight of planet Mercury

1 FEBRUARY

Jupiter, Venus, Saturn and an 11% waning crescent Moon visible in the morning sky

Mars and Uranus appear 1° apart

14 FEBRUARY

Jewelled Handle effect visible on the Moon

19 FEBRUARY

Closest and largest full Moon of 2019

Callisto sits just north of Jupiter's disc as dawn breaks



Lebruar at a glance

February is an interesting month for UK astronomy. Although the nights are starting to get shorter, the rate of shortening isn't that fast, especially at the start of the month. In 2019 this conveniently coincides with a new Moon

The bright stars of winter are still very evident but also notably drifting ever further towards the west. Following from the east are the more subtle constellations of spring, led by the very recognisable form of Leo the Lion. This is one of the few constellations which actually bears a good resemblance to what it's supposed to represent. The Lion's head is marked by a distinctive asterism known as the Sickle. Resembling a backwards question mark, the punctuation dot is provided by the bright star Regulus (Alpha (α) Leonis). The second star in the Sickle above Regulus is Algieba (Gamma (y) Leonis), a close binary. Don't be fooled by mag. +4.8 40 Leonis 22.5 arcminutes south of Alaieba, as the actual binary is much tighter, separated by just four arcseconds. Leo is one of 13 constellations through which the Sun passes during its annual journey against the background stars. Immediately to the west of the

MOON PHASES

Key stages in the monthly cycle









VISIBLE PLANETS

Where to spot the planets this month





MERCLIRY Evening object from mid-month that reaches great elongation on 27 February.

VENUS Morning planet, near to a waning crescent Moon and Jupiter at the

Evening object that appears small through a telescope.

start of February.

A morning planet found close to Venus and a waning crescent

Moon on 1 February

A morning planet that appears 2 February. Occultation and southeast England.

URANUS Evening planet, best seen at the start of the month.

NEPTUNE Dim Neptune is rapidly swallowed by the evening twilight during

west and you'll arrivGemini, the Twins. This is Ħ

ite an easy constellation to recognise thanks to the two bright 'twin' stars, Castor (Alpha (a) Geminorum) and Pollux (Beta (6) G eminorum). These represent the heads of Castor and Pollux, the half-twin brothers from Greek and Roman mythology, with the rest of the constellation providing a stick-figure representation of their bodies, arms and leas, Although the two stars appear very similar at first glance, look carefully and you sh ould be able to see that they are q uite different. Pollux is slightly brighter and more orange in colour.

A beehive and a snake

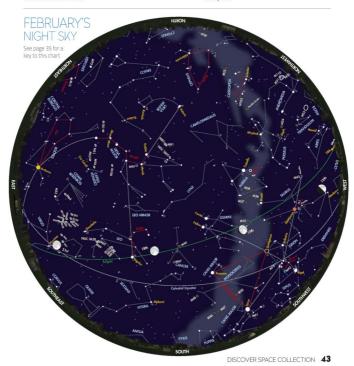
Due south of Pollux is another bright star which looks rather isolated in the sky. This is Procyon (Alpha (a) Canis Minoris). If you join the dots by drawing an imaginary line between Procyon, Sirius (Alpha (a) Canis Majoris) to the southwest and orange Betelgeuse (Alpha (a) Orionis) in the northeast corner of Orion, this creates a large asterism known as the Winter Triangle. Returning to the northern twin star Castor, draw an

imaginary line from it to Regulus at the foot of the

▲ Locate Mars in view too

Sickle Slightly south of the mid-way point along this line is a fuzzy patch of sky just viable to the naked eye under clear conditions. This is the Beehive Cluster. M44, which sits at the heart of Cancer. The Beehive is best seen with low-power optics such as binoculars or a telescope using a low magnification. It's interesting to compare the appearance of M44 with another pen cluster in Cancer, M67, which lies near to Acubens (Alpha (a) Cancri). This is an ancient object estimated to be 5x further away than M44. This is most evident when you view M67 through a telescope, as it appears much smaller and dimmer.

Cancer looks like a faint inverted-Y shope in the skyl. If you look south of the open end of the Y, you addid find a distinctive sideways tear-drop pattern of stars representing the head of Hydra the Watersnake. This is the largest constellation to prace in the entire night sky and takes a staggering nine hours to completely itse above the UKs horizon. The star Alphard (Alpha (a) Hydrae) marks the snake's neck and can be seen as the only medium bright star in a rather barren area of sky southeast of Hydra's head. Orange Alphard's name rather appropriately translates as 'the solitary one'.



The Triangulum Galaxy ▷

Alec Alden, Colchester, 17, 18, 24 and 25 November 2017

HOTSHOTS WINNER MARCH 2018

March

A chance to enjoy galaxyladen Leo as the spring equinox arrives

1 MARCH

Venus, Saturn, Jupiter and a 23% waning lunar crescent line up in the morning sky

13 MARCH

Lunar X and V effects on the Moon visible at 16:43 UT

18 MARCH

Ganymede shadow transit from 04:00 UT

27 MARCH

Jupiter 1.1° from a 60% waning gibbous Moon at 01:00 UT

31 MARCH

Mars 3.2° south of the Pleiades this evening



▲ Catch the clair obscur effects known as the Lunar X and V in daylight on the afternoon of 13 March. The effects are at their peak around 16:45 UT

March

The March equinox occurs at 21:58 UT on 20 March and indicates the point in time when the centre of the Sun's disc crosses the projection of the Earth's equator into the sky, a great circle known as the celestial equator. At this equinox, the crossing marks the Sun's transition from the southern to the northern half of the sky. The equinoxes also mark the time when the length of night and day is changing at its fastest rate. The March equinox, also known as the northern hemisphere's spring equinox, marks the time when the rate of shortening of the night is at its greatest. We'll have to wait six months for the September equinox, when the rate of nighttime lengthening is at its greatest. All this serves as a reminder to get a last look at magnificent Orion before he gets swallowed up by the advancina evening twilight! Leo, the Lion is the most

prominent constellation due south for much of the month. The Lion's head is depicted by the Sickle asterism with the creature's body indicated as a rectangular shape extending to the east. The mediumMOON PHASES Key stages in the monthly cycle









VISIBLE PLANETS

Where to spot the planets this month





MERCLIRY Well placed evening planet at the start of March, but lost from view

from 10 March VENUS

A brilliant morning object close

MARS An evening planet passing south of the Pleiades at the end

A morning planet that moves into a favourable, albeit low,

position towards the end of the month.

SATURN Morning planet. Saturn has a close encounter with a 40% vaning crescent Moon on 29 March

URANUS An evening planet that is lost from view towards the end of the month.

NEPTUNE Not visible

bright star Denebola (Beta (ß) Leonis) marks the end of the Lion's tail. The depiction of the cat is completed by four legs hanging below its body. The rear legs are convenient for locating a number of galaxies visible in the region.

Galaxies galore

The most famous galaxies in Leo are collectively known as the Leo Triplet, these being M65, M66 and NGC 3628. They are visible through binoculars

ith more definition through a small telescope. a nd are interesting because they exhibit n oticeably different shapes. Located near th e west-pointing front leg is another g roup of galaxies, known as than 96 Group. These are elliptical galaxies

with little discernible structure. which appear as fuzzy patches when seen through a small telescope. The M96 Group includes M95, M96 and M105,

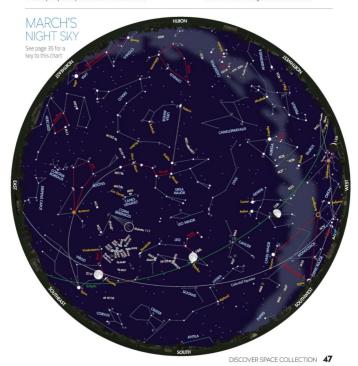
To the east of Leo is a lovely triangular smattering of faint stars th at appear to shimmer in and out of visibility on a clear, dark March night. This is the open cluster Melotte 111, representing t he beautiful sweeping locks of Queen Berenice

▼ A 60% lit

and forming part of the constellation Coma Berenices. The cascade of stars appears to flow south into the open Bowl of Virgo, a name describing a large asterism of stars in the constellation of Virgo, resembling a semi-circle.

The region between Melotte 111 and the Bowl of Virgo is littered with fairty galacies, many of which can be seen with a small telescope. The region is so rich in galactic fare that it has gathered the name the Realm of Galacies. The reason why there are so many galacies on view here is twofold. First, the plane of our wan Milliy Way Galacy which is so evident in summer and winter rotates out of the way during spring and autumn allowing us to peer out into the deepest depths of space. Secondly, there are two large galaxy clusters in this direction: the Coma and Virgo galaxy clusters.

Galaxies remain important as you head north out of Coma Berenices and into **Canes Yenatici**. Although a small constellation containing just two prominent stars, it, plays host to Md3 the Sunflower Galaxy, M64 the Blackeye Galaxy, M51 the Whitppool Galaxy, M64 the Blackeye Galaxy, M51 the Whitppool Galaxy and whole host of other, less well-known objects of similar ilk. Many of the galaxies mentioned are bright enough to be seen with nothing more than binoculars.





Constellation of the season

WINTER: Orion

The stars of Orion represent a mighty hunter of chequered reputs, whose eventual undoing came during a hunting trip on Crete where, occompanied by Artemis, goddess of hunting, he boastfully threatened to till every living creature on Earth. An angry Mother Earth sent a giant scorpion to teach Orion a fatal lesson a Cause placed Orion and the scorpion in the heavens as memorial. Understandably, Scorpius and Orion can be found on a possible sides of the skill.

Orion's Belt is formed from three stars of similar bightness, equally spaced in a limb Mintaka, Ainlam and Aintak. The western star, Mintaka, is positioned 17 arcminutes south of the celestial equator and effectively traces the path of this great circle throughout the course of a night. The open cluster Collinder 70 has an apparent diameter around 3" and includes the three Belt stars along with Sigma Orionis. The cluster is 800–900 lightyears distant and contains around 190 members bighter than 1014 magnitude. Bincoulars provide the best view. Orion's shoulders' are marked by red superciant stars bestleaves and Bellatrix.

Treasures among the clouds
The main stars below Orion's Belt are Rigel to the
southwest and Saiph in the southesst. Blue supergiant
Rigel represents Orion's foot and is the seventh brightest
star in the night sky. It's around 80 times larger than
the Sun and more than 172,000 times more luminous.
Moving your gaze between Rigel and Betelgeuse really
bings out their colour difference.

The three Belt stars, Betelgeuse and Rigel are all

located in a minor spiral arm of the Millicy Way known as the Orion Spur. This diso plays host to the Orion Nebula, M42, a glowing gas cloud in the centre of Orion's Sword which is represented by a line of faint stars below his Belt. A cluster of stars known as the Trapezium have formed from the nebula material and are currently responsible for ionising the surrounding gas. cousing the rebula to glow. Immediately north of M42 is comma-shaped M43, another clowing nebula.

Orion plays host to the Horsehead Nebula, B33. This is formed by a dark finger of nebulosity silhouetted against a bright curtain of glowing hydrogen known as IC 434. The silhouette resembles the profile of a horse's head. The Horsehead is located south of Alnitak and, although not easy to see visually, is a very rewarding photographic target. A brighter and visually easier nebula sits east of Alnitak the Rome Nebula, NGC 2024. Another noteworthy object not too far from the Belt is the beautiful reflection nebula NF3. Which lies close to another reflection nebula (NF2.071.

Long-exposure images reveal an immense ring of glowig-exposure images reveal an immense ring of glowing red hydrogen in Orion. Known as Barnard's Loop, the ring measures 10° across. Further north, long exposures also reveal the smaller, circular nebula centred on Meissa. This is known as the Angelfish Nebula. Sharaless 2264.

Although Orion is brightest in winter, look out for the Orionid meteor shower from 2 October until 7 November. Associated with Halley's Comet, it delivers a peak ZHR of around 20 meteors per hour, appearing to come from the region depicting the hunter's club.











Martin Baker, 25 November 2017 & 18 January 2018, Berkshire

HOTSHOTS WINNER APRIL 2018

April

Fingers crossed for the zodiacal light on Moonless evenings this month

START OF APRIL

The zodiacal light may be visible under dark skies

Mars, the Pleiades and Hyades group together

14 APRII

Jewelled Handle effect visible on the Moon

19 APRIL

Favourable lunar libration for the southeast limb

23 APRII

Bright Moon and Jupiter close in the morning sky

Orll at a glance

The spring sky is in full force during April. Directly overhead is that most recognisable of asterisms, the Plough or Saucepan, its seven stars creating a familiar shape that, from the UK at least, never sets below the horizon. This makes it very useful for navigation. Viewing as a saucepan, locate the two stars furthest from the handle. These are Dubhe (Alpha (α) Ursae Majoris) and Merak (Beta (β) Ursae Majoris), together known as the Pointers because extending their line to the north will bring you to the Pole Star, Polaris.

Polaris is a medium-bright star around which all others appear to rotate as seen from the Northern Hemisphere. If you were at the North Pole. Polaris would appear directly overhead at night. Once you've found Polaris, dropping a vertical line from it to the horizon identifies due north. In addition, the altitude of Polaris above the northern horizon equals your latitude. It's a very useful star indeed

Getting a handle on M3

Return to the Saucepan and extend the arc of its handle away from the pan to eventually reach the bright orange giant Arcturus (Alpha (a) Bootis). Keep the arc going to

MOON PHASES

Key stages in the monthly cycle







5 Apr 12 Apr

VISIBLE PLANETS

Where to spot the planets this month





MERCURY Not visible

net. An 84% lit waning on 23 April.

na planet best seen Bright me at the start of the month.

SATURN Morning object not far from nearby on 25 April.

Pleiades and Hyades at the start

of April. Less visible by month end.

URANUS Not visible.

HIDITED Dominant but low morning

NEPTLINE Not visible

reach brilliant white Spica (Alpha (α) Virginis). Visualising the Saucepan's handle arc as part of a circle, at its centre is Cor Caroli in Canes Venatici. Draw a line from Arcturus to Cor Caroli and at the mid-point sits M3, a superb globular cluster to view through any size of telescope.

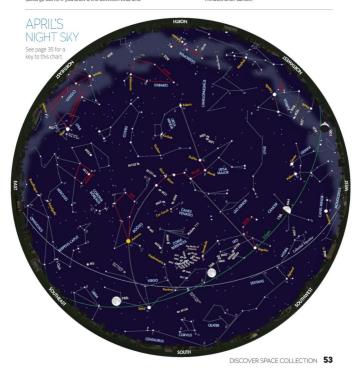


▼ All eyes to

Spica is the brightest star in Virgo the Virgin, the second largest constellation in the night sky. The largest is **Hydra the Matersnake** which slithers south of Virgo. Between the western region of Virgo and the Snakes body sit two small constellations known as Crater the Cup and Convus the Crow. Crater has a faint but distinctive shape but requires dark skies to see properly. Corvus is smaller and brighter. It contains an asterism known as the Soil because being of low altitude it resembles as alling boat's sail diffiting along the horizon.

The star in the southeast corner of the Sail is Kraz (Beta (6) Corvi). If you draw a line between Kraz and Porrima (Gamma (n) Virginis), the star at the bottom of the large asterism known as the Bowl of Virgo, at the mid-point is the small but distinctively shaped M104, the Sombrero Galaxy. Larger instruments will detect a dark dust lane passing across its longest dimension.

A new Moon will leave the start of April good and dark - a perfect time to look to the vess of ther sunset to try to see the dim conical glow of the zodiacal light. This represents sunlight scottered by dust in the plane of the 50xiar System. You'll need to find a dark site to make your ottempt. Typically, the best time to start looking is 90 minutes after sunset.



May

Dark skies could make for exciting Aquariids and M5 is well placed mid-month

Eta Aquariid meteor shower reaches its peak this morning

Lunar X and V will be visible at 17:45 UT

20 MAY

Moon close to Jupiter as darkness falls

26 MAY Approximate start of

the noctilucent cloud season

28 MAY

Ceres reaches opposition on border

Comet C/2016 R2.
(PANSTARRS) & the Pleiades

osé J Chambó, New Mexico, US, 4 February 2018

DISCOVER SPACE COLLECTION 55.

Vat a glance

MOON PHASES Key stages in the monthly cycle







It's during May that the nights really appear noticeably shorter. Certainly by the end of the month, it can seem that no sooner have you set up for night observing than dawn twilight starts to appear, an effect that's worse the further north you are in the UK. Despite this, there are still plenty of things to look out for.

The bright orange star Arcturus (Alpha (α) Boötis) is very evident indeed. It sits at the southern 'pointed' end of a large kite-shaped asterism. Being as high as the Saucepan is long, the Kite represents the main body of Boötes the Herdsman. For such a large constellation it's somewhat surprisingly devoid of deep-sky objects, but does contain a number of impressive double stars. Turn to page 62 for more details on these.

Running between Boötes and Ursa Minor, the Little Bear, is the winding constellation of Draco the Dragon. The Dragon's head is marked by a distinctive quadrilateral pattern of stars known as the Lozenge.

East of Boötes is the lovely compact constellation of Corona Borealis, the Northern Crown, This is a small semi-circle of stars often described as a backward C. Its brightest star is Gemma (Alpha (a) Coronae Borealis), the 'jewel' in the crown. The inside of the semi-circular crown is, to the naked eve, devoid of stars unless you're lucky enough to catch an irregular variable known as R Coronae Borealis at peak brightness. This remarkable star is carbon rich. Sooty deposits build around it, blocking its light so it dims and appears quite red. Over time the soot clears, allowing R to increase in brightness once again. This has earned R the informal name the 'Dust Puff Star'. If it is visible, it'll be right on the edge of naked eve visibility from a dark-sky site.

Southeast of Corona Borealis is the zig-zag shape of Serpens the Serpent Serpens is unique in the heavens in that it's split in two. The head is represented by Serpens Caput and the

tail by Serpens

▼ The Eta Aquariid lar asterism durina the shower's activity period. Expect peak rates

VISIBLE PLANETS Where to spot the planets this month



MERCLIRY Morning planet for most of May. but best in evening sky at the end of the month

Bright morning planet, poorly positioned. Best seen at the end

Evening planet compromised by northeast edge of M35 on 19 May. Bright object in southern Ophiuchus. Low from the UK, only reaching altitude of 15°.

Morning object in Sagittarius Bright 82% waning gibbous Moon 2.1" southeast of Saturn on 23 May.

Not visible NEPTUNE

Cauda, Serpens is being carried across the sky by the constellation Ophiuchus the Serpent Bearer. It is Serpens Caput that lies west of Ophiuchus and southeast of Coronae Borealis. The magnificent globular cluster M5 is located between Serpens Caput and the northeast corner of Virgo. It's conveniently located very close to the faint star 5 Serpentis. Find this star and you'll find M5.

Tracking magnificent M5

One way to locate M5 is to use the constellation of Librathe Scales which lies to the south. The two brightest stars in Libra - Zubenelgenubi (Alpha (α)

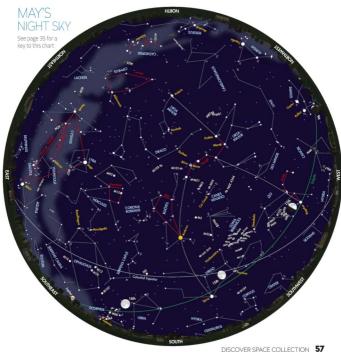
Librae) and Zubeneschamali (Beta (B) Librae) - form what's best described as a 45° line in the sky. Imagine Zubeneschamali as the centre of a threepronged steering wheel, with Zubenelgenubi as the star marking the

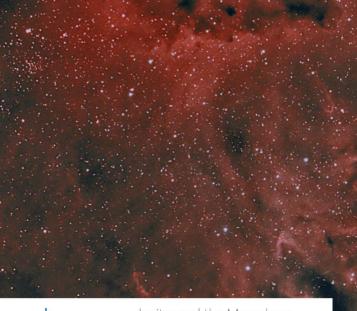




end of the bottom-left spoke. M5 would be just to the north of where you'd imagine the top spoke to end.

The dark moonless sides at the start of May are ideal for watching the peak of the Eta Aquartial meteor shower, which has a ZHR of 40 meteors per hour at peak on the morning of 6 May. Daylight doesn't need to halt observing either, and the lighting effects known as the Lunar X and V reach their peak in the afternoon of 11 May, when a giant letter X may be seen using a telescope approximately 1/3rd the way up the terminator from the south point at 17x45 UT. The V is about 3/5ths of the way up the terminator.





Jupiter and the Moon Keep company as giants Hercules and Ophiucus dominate Jupiter and the Moon keep

ALL MONTH Noctilucent cloudspotting season

8 JUNE Europa and its shadow transit virtually aligned

10 JUNE Jupiter reaches opposition

15 JUNE Almost full Moon. Jupiter and Antares are nicely grouped

28 JUNE Peak of the June Boötids, with low activity expected

DISCOVER SPACE COLLECTION 59



at a glance

June is a testing month for UK astronomy because the nights are so short. A number of less distinct constellations try to take centre stage, but close on their heels are the strong patterns of summer. The view to the south is dominated, if that's the correct term, by two giants of the night sky: Hercules the Strongman and Ophiuchus the Serpent Bearer, Hercules is a faint constellation but redeemed by an asterism known as the Keystone. This is formed from Epsilon (s). Zeta (C). Eta (η) and Pi (π) Herculis.

Although the Keystone is not particularly bright, its shape does stand out rather well and acts as a guide for finding the Great Globular in Hercules, M13. This globular cluster sits one-third along the side of the Keystone from Eta toward Zeta Hercules, It's a really magnificent example, containing around 100,000 to 1 million stars, packed into a spherical volume just 145 lightvears in diameter.

The head of Hercules is located at the southern extremity of the constellation, marked by the star Rasalgethi (Alpha (a) Herculis). Hercules is upside down in UK skies, his head located close to the head of our second giant, Ophiuchus. The Serpent Bearer's head is represented by Rasalhague (Alpha (α) Ophiuchi). From there the main body of the constellation appears to

MOON PHASES

Key stages in the monthly cycle









VISIBLE PLANETS Where to spot the planets this month



MERCURY elongation of 25.2° on 23 June. Near Mars on 18 June

VENUS Morning object rising 45 minutes before the Sun all month.

MARS

▼ The asterism

Star and open

Poorly positioned evening planet, appearing close to Mercury on the evenings of 17-19 June.

Reaches apposition on 10 June Rather low as seen from the UK.

Morning planet visible to the east of the Teapot in Sagittarius.

Not visible MEDTLINE

Tricky morning planet. Possibly visible close to Phi (a) Aquari

the south as a box shape that some have likened to an upturned flowerpot. Two leas dangle from the southern corners of the box, with bright Jupiter currently wandering through this area.

Ophiuchus contains a number of interesting globulars, including M10, M12 and M19. His eastern shoulder plays host to IC 4665, a large and relatively bright open cluster 1.3" north of Cebalrai (Beta (B) Ophiuchi). Look 4.5" east of Cebalrai and, given clear skies, you should be able to see a faint V pattern of stars. This is known as Poniatowski's Bull. Barnard's Star sits west-northwest of the star, marking the end of the western arm of the V (66 Ophiuchi). Barnard's Star has the largest measured

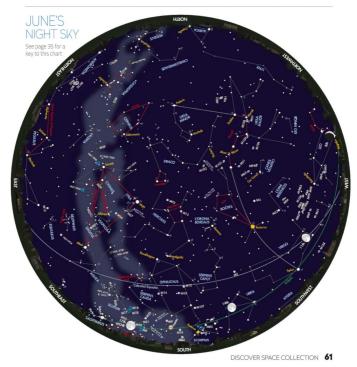
Barnard's Star IC 4665 **OPHIUCHUS**

proper motion in the entire sky, moving roughly half the apparent diameter of the Moon against the fixed background sky every 90 years. At mag. #9.5, a telescope is required to see Barnard's Star.

Belov Ophiuchus's western leg is bright orange Antares (Alpha (a) Scopii), the star marking the heart of Scoppius the Scoppion. The binocular globular cluster, Ma sits 1.3' west of Antares. The region east of the star tokes us into a very rich part of the Milky Way – the plane of our own home Galaxy, represented by a mistry path formed from the integrated light of billions of disknot stars. The region immediately east

of Antares is the part of the Milky Way we see marking the direction of the bright core of our Galaxy.

In the same general region as Antares is Jupiter, which on 10 June is at opposition (in the apposite part of the sky to the Sun). On the evening of 16 June the Moon is also at opposition, and can be seen very close to Jupiter. June is also a good time to see noctificent clouds. Look out 90–120 minutes after sunset, low dove the northwest horizon, or a similar time before sunrise, low above the northwest horizon, or a fine for a proper with an electric blue colour, often with a fine wispy structure glowing against the dank of night.





Constellation of the season

SPRING: Boötes

The constellation Boötes the Herdsman has various myths associated with it. In one it represents Arcas, the illegitimate son of Zeus and Callisto, Angered by Zeus's infidelity, his wife Hera turned Callisto into a bear, When Arcas came face-to-face with his mother in the woods, he drew an arrow to shoot her before Zeus intervened and turned Arcas into a bear. Grabbina mother and son. he swung them with such ferocity their tails stretched and they ended up in the heavens as Ursa Major and Ursa Minor, the Great and Little Bears. The Greek name for Boötes is Arctophylax, which means 'Bear Watcher'.

Boötes is a large kite-shaped constellation. Its primary star is Arcturus (Alpha (a) Boötis), the fourth brightest nighttime star, which shines with a distinct orange colour. Arcturus has the largest proper motion of all the first-magnitude stars, with the exception of Alpha Centauri. It appears to move 2.29 arcseconds a year relative to the stellar background.

The asterism that forms the main body of the herdsman is 23.5° high, extending north-northeast from Arcturus. Another, considerably smaller and fainter, small telescope asterism, Picot 1 can be found half a degree south-southwest of Arcturus and is formed from seven stars around 10th magnitude that depict the profile of Napolean's hat!

There is a surprising lack of deep-sky objects in Boötes. The most conspicuous include the mag. +10.5 globular cluster NGC 5466, the mag. +10.2 spiral galaxy NGC 5248 and the mag. +10.9 spiral galaxy NGC 5676. However, there's an empty expanse of space in this

direction known enigmatically as the Boötes Void. This roughly spherical region of the Universe measures 330 million lightvears in diameter and contains hardly any galaxies. The centre of the void lies 700 million lightyears from the Sun.

Arcturus's beautiful companions

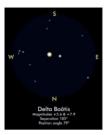
One thing Boötes does excel at is multiple stars, many of which are easy to split with binoculars or a small telescope. Highlights here include Delta (δ) Boötis which is very easy as a mag. +3.6/+7.9 pair separated by 103.8 arcseconds. Kappa (κ) Boötis is harder, with magnitude +4.6/+6.6 components separated by 13.5 arcseconds. A low-power eyepiece also allows lota (ı) Boötis to be viewed in the same field as Kappa, with mag. +4.8/+7.4 components separated by 38.7 arcseconds.

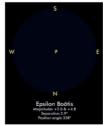
Pi (π) Boötis is a lovely tight double with component magnitudes of +4.9/+5.8 separated by 5.5 arcseconds. Xi (E) Boötis is another worthy target, showing as a yellow and orange pairing of mag. +4.8/+7.0 stars separated by 6.3 arcseconds

Izar (Epsilon (s) Boötis) shows an unequal mag. +2.6/ +4.8 pair separated by a narrow 2.9 arcseconds. You'll need a 75mm or larger scope to split Izar, but the effort is well worth it as they shine with golden-yellow and blue colours.

Boötes plays host to several annual meteor shower radiants. The most famous are the Quadrantids, with a peak on 3/4 January with a ZHR of 120 meteors per hour and favoured by a new Moon on 6 January.

▼ What Boötes lacks in deepit makes up for









July

The nights may be short but Saturn's rings shine brighter than normal this month

ALL MONTH
Noctilucent clouds

Noctilucent clouds

9 JULY Saturn reaches opposition 14 JULY

Pluto reaches opposition, 3° north of 52 Sagittarii

16 JULY Partial lunar eclipse 29 JULY

Peak of the Southern Delta Aquariid meteor shower





JUIV_{at a glance}

The bright, vibrant constellations of summer now take their place centre stage towards the south. The most prominent pattern on view is that of the giant asterism known as the Summer Triangle. Formed from the three brightest stars in Cygnus, Lyra and Aquila - respectively Deneb (Alpha (α) Cygni), Vega (Alpha (α) Lyrae) and Altair (Alpha (α) Aquilae) - the Triangle plays host to a bright section of the summer Milky Way.

Imagining the Summer Triangle to be pointing down towards the horizon, slightly west of the region it's pointing at sits another distinctive pattern of summer - the Teapot asterism. This is part of Sagittarius the Archer and really does look like its name suggests. The spout off to the west marks the approximate position of the centre of our Milky Way Galaxy.

It's a very worthwhile exercise scanning the region

where you'd expect steam to be rising out of the Teapot's spout. Here you'll find exotic objects such as M8 the Lagoon Nebula and M20 the Trifid Nebula. Both sit well within a binocular view, providing a great opportunity to compare their appearance. The

Altitude 13



▼ Circumstances

20:00 UT as seen

from the centre

MOON PHASES

Key stages in the monthly cycle







VISIBLE PLANETS

Where to spot the planets this month





MERCURY Not particularly well placed

VENUS Bright morning planet, but close to the Sun. Moon near on the

MARS Evening planet, lost from view

HIPITER Bright planet, well-placed but low as seen from the UK.

Reaches apposition on 9 July Rings brighten leading up to opposition, fading in the days after.

URANUS Uranus is a morning planet, best seen towards the end

NEPTUNE Neptune is a morning planet. Its position improves throughout the month.

Lagoon is expansive and bright, while the more distant Trifid appears dim and compact. Additional rewards in the area include the rich open cluster M25 and equally resplendent globular cluster M22.

North of the Teapot is multi-titled M17, also known as the Swan, Lobster, Omega and Horseshoe Nebula. Between M17 and M20 you'll find the large rich starcloud known as M24 or Delle Caustiche.

North of M17 is M16 the Eggle Nebula. This can be a frustrating object, showing as a beautiful glowing gas cloud through long-exposure photography. It's here that the shape of a spread-winged eagle appears. However, visually through the eyepiece the nebula is often missing, leaving only the star cluster at the heart of M16. Larger apertures will show it, but it's much fainter than

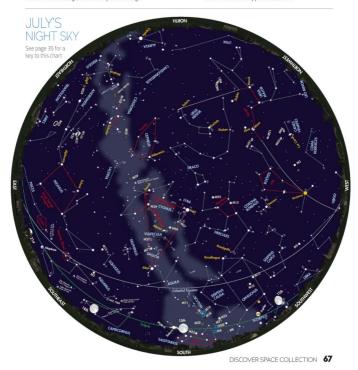
photographs would suggest. M16 is technically within the constellation of Serpens Cauda, the Serpent's Tail, lying close to the border with Scutum the Shield. This is another rich deep-sky region thanks again to the presence

of the Milky Way, Scutum looks like a narrow diamond of stars oriented northeast/scutiwest. Its northern extreme lies close to the 'tail' of **Aquila**, a region marked by a gentle curve of faint stars. The curve conveniently runs towards Mtl the Wild **Duck Cluster**, one of the richest and most impressive open clusters in the entire injalt sky.

The planet **Saturn** is currently located in Sagittarius, just south of another asterism known as the Teaspoon, itself located northeast of the Teapot. Saturn reaches opposition on 9 July and in the days running up to this date it's worth taking a look at the planet through

a telescope. The rings brighten in the days before opposition and fade back to normal in the days that follow, an optical phenomenon known as the Opposition Effect.

Six months on from the total lunar eclipse at the start of the year, another **lunar eclipse** is set to occur on 16 July. Unlike the one in the early hours of 21 January, this month's eclipse will be a partial, just beginning as the Sun sets and the Moon rises on the evening of 16 July. Maximum eclipse occurs at 2131 UT when the Earth's shadow will encroach to cover 65 per cent of the Moon's apparent Gimmeter.



Centaurus A >

Fernando Oliveira De Menezes. São Paulo, Brazil, 18 May 2018

HOTSHOTS WINNER AUGUST 2018

August

All eyes on Lyra, Cygnus, Aquila and bright Jupiter for Ganymede's shadow transit

1 AUGUST

Mars lies just south of an ultrathin lunar crescent after sunset

Perseid meteor shower peaks under bright Moonlight

22 AUGUST

Ganymede's shadow transits Jupiter from 19:18 UT

26/27 AUGUST

Jupiter passes across globular cluster NGC 6235 in Ophiuchus

30 AUGUST

Thin morning Moon with Mercury nearby

▲ There are two thin Moon-spotting opportunities this month, the one on 30 August,

August is an exciting month. As the nights start to lengthen, the sky views on offer are nothing short of breathtaking. Stars and constellations which were so tantalisingly on offer during July's short nights slowly drift to the west as the retreating evening twilight gives us longer to look at them.

The Summer Triangle is now very dominant high in the south. Its vertex stars, Deneb, Vega and Altair are very obvious. Vega (Alpha (α) Lyrae) in the northwest is the visually brightest of the three. For a long time this was the star that marked the zero point on the magnitude scale used to describe star and object brightness. Its host constellation is Lyra the Harp. This is a compact affair best defined by the squashed diamond pattern of fainter stars that 'hanas' south of Vega. The two stars at the base of the pattern. Sheliak (Beta (B) Lyrae) and Sulafat (Gamma (v) Lyrae), provide a convenient method for locating a showpiece of the summer skies, the Ring Nebula M57. This object, representing the outer layers of a low-mass dying star ejected off into space, sits slightly offset towards # Sheliak from the mid-point of the line between both

MOON PHASES

Key stages in the monthly cycle



1 Aug







VISIBLE PLANETS

Where to spot the planets this month





Morning object at greatest estern elongation on 9 August. Lost from view after 24 August.

VENUS Not visible

▼ Through

can be seen

Near a very thin waxing crescent Moon on 1 August, but soon becomes lost from view.

JUPITER Low evening planet. A 72% lit

axing crescent Moon lies close on 9 August.

SATURN Evening planet in Sagittarius, positioned under the Teaspoon.

LIRANUS

Improving morning planet on the verge of naked eve visibility.

Morning planet best seen towards the end of the month. close to Phi (¢) Aquarii.

stars. It's also a fraction south of the line. Through a telescope, M57 is easily overlooked as a faint star at low magnification, but as the power is piled on, its extended disc is most evident.

Seeing doubles in Lyra

Vega marks one point of an equilateral triangle, with Zeta (E) and Epsilon (E) marking the other vertices. Look hard at Epsilon to the north and see if you can split it into two stars. Through a telescope the double nature is very evident, but there's a further twist because each of the components is double again. The extra doubling can be hard to see at first, good optics and a reasonable magnification being required. This physical



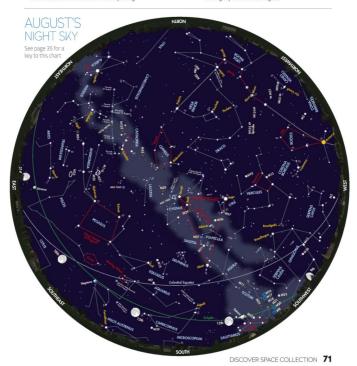
aspect of the Epsilon Lyrae system has given it the name the Double-Double.

As we're describing doubles, one of the most beautiful in the entire sky lies to the southern end of a large asterism known as the Northern Cross.

Albirea (Beta (8) Cygni) is a glorious sight through a telescope, presenting a golden-yellow primary and azure blue secondary, it marks the head of Cygnus, a constellation that represents a swan flying across the night sky. It's an easily identified pattern rich in deep sky wonders. Turn to page 76 to find out more about it. This month there are two thin Moon-soutine.

opportunities, a tricky one after sunset on 1 August and one before sunrise on 30 August. Both have planets nearby, Mars on the 1st and Mercury on the 30th, which should make finding these thin slivers a little easier.

Another planetary event worth noting is the apparent passage of Jupiter in front of the globular cluster NGC 6235 in Ophiuchus. The globular has an apparent size of 5 arcminutes, eight times larger than the apparent size of the planet. Jupiter's bightness will be an issue, but astrophotography may be able to reveal this unusual conjunction visible in the early evening sky of 26 and 27 August.







September

Plenty of evening viewing this month, and Neptune skips past red giant Phi Aquarii

Mag. +7.8 Neptune is 30 arcseconds from star Phi Aauarii

10 SEPTEMBER

Neptune reaches opposition

Mag. -0.9 Mercury is just 18.6 arcminutes from -38 Venus after sunset

14 SEPTEMBER

Harvest Moon

The northern hemisphere's autumn equinox occurs at 07:50 UT

September at a glance

MOON PHASES

Key stages in the monthly cycle





22 Sep







14 Sep **VISIBLE PLANETS**

Where to spot the planets this month





After superior conjunction on 4 September, emerges into a

poor position in the evening sky.

Evening object close to the Sun. A 1% waxing Moon is 3.3° above Venus on 29 September

Not visible

Bright Jupiter is an evening

Evening object in Sogittarius. A 76% lit waxing gibbous Moon is nearby on 8 September.

URANUS Currently the best positioned planet, reaching 50° altitude in darkness when due south.

NEPTLINE Neptune reaches opposition on 10 September. Very close to Phi (a) Aquarii on 6 September.

The nights really start to draw in quickly during September, an effect caused by the apparent southward motion of the Sun reaching a peak rate for the year. The peak occurs at the September equinox. also known as the northern hemisphere's autumn equinox. This is an instant when the centre of the Sun's disc crosses the projection of Earth's equator into the sky, a great circle known as the celestial equator.

There's a certain excitement to the sky at this time of year because dark nights are back on the cards. The stars of summer remain well placed, with the Summer Triangle still in a dominant southerly position early evening. The integrated light of billions of stars in our own Galaxy forms the misty path known as the Milky Way and this can be seen, from a dark-sky location. flowing down through the triangle. A detached patch of Milky Way close to the bottom of the constellation of Aquila the Eagle is renowned for tricking sky watchers into thinking clouds are starting to appear. Known as the Scutum Star Cloud, this bright, heart-shaped patch of Milky Way is very evident from a dark-sky site.

to mag. +4.2 Phi Aquarii on the

▲ Neptune

A fox, an arrow and a dolphin The region to the north of Aquila is populated by

small constellations, some well-defined, others frankly obscure. At the foot of the Northern Cross, near to the be autiful binary star Albireo (Bet#)(Cygni we st art at the obscure end of the spectrum

w ith the very faint and rather loose form of Vulpecula the Fox. Despite its illd efined nature, there are a number

of interesting objects to view here. Brocchi's Cluster Collinder 399 is located in the southern part of the constellation, 4.5° south of Alpha (a) Vulpeculae. The stars here form the shape of a coat hanger, giving the object the informal name t he Coathanger Cluster. You can just ab out see the coat hanger shape with

t he naked eye, but you're not looking a t a true cluster here as Collinder 399 is an a sterism, a chance alianment of stars,

The beautiful planetary nebula M27, the Dumbbell Nebula, appears to the east-northeast of Collinder 399. Visible in binoculars, this object takes on the appearance of an apple core when seen through the eyepiece of a telescope. A good way to locate it is to use another small constellation found south of

▲ The full Moon

The brightest star in Aguila is Altair (Alpha (a)

Aquilae). It marks the bottom vertex of the Summer

Triangle and is easy to identify because it has two

and Tarazed (Gamma (y) Aquilae).

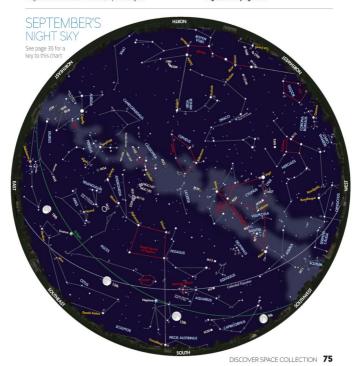
dimmer stars either side of it: Alshain (Beta (6) Aquilae)

Vulpeculo. Segitta the Arrow has better definition than Vulpeculo and is cartually shaped like an arrow. Locate the star Delta (a) Sagittae which marks the point where the shaft meets the flight of the arrow. Imagine the line from Delta to Gamma (g) Sagittae, which is at the pointed end of the arrow's shaft. Rottet he line for 120 'around Gamma in a counter-clockwise fashion. Where Delta would end up marks the position of M27.

East-southeast of Sagitta is the small and distinctive constellation of **Delphinus the Dolphin**. It looks like a tight diamond of stars with a tail, representing the

nose and neck of a bottlenose dolphin. An interesting fact about Delphinus is that it contains the star Rho (g) Aquilae. Since this star was labelled as part of the constellation of Aquila, its appearent motion in the sky has made it drift slightly across the border between Aquila and Delphinus, so it's now in the wrong constellation!

The line of small constellations ends as it began, with the small and rather obscure constellation of **Equuleus the Foal**. Its four faint stars nestle up against another equine pattern, the autumn constellation of **Peaasus the Flying Horse**.





Constellation of the season

SUMMER: Cygnus

Cygnus the Swan flies high above UK skies during the summer. Its central pattern, the large and easily identified asterism known as the Northern Cross. represents the tail, body and wing-stubs of the Swan.

Myth tells of Cycnus and Phaethon, the son of the Sun god Helios, racing across the heavens. Venturing too close to the Sun, their chariots burned up and fell to Earth, Phaethon fell in the river Eridanus and drowned. Cycnus survived and asked Zeus to turn him into a swan so he could retrieve his friend. Zeus was so moved by this that he placed Cycnus in the heavens.

Bright Deneb (Alpha (a) Cygni) sits on top of the cross with dimmer Albireo (Beta (6) Cvani) at the base. Deneb represents the bird's tail, and Albireo, a beautiful binary, its beak, Although Deneb appears outshone by the other two triangle stars, Vega (Alpha (a) Lyrae) and Altair (Alpha (a) Aquilae), that's because they are only 25 and 17 lightyears away. At an estimated 2,500 lightyears distance. Deneb is far brighter, ground 200,000 times more luminous than the Sun and 200 times larger.

Riches of the deep sky

The presence of the Milky Way gives Cygnus a rich repository of deep-sky objects. Long exposures reveal extensive glowing hydrogen in the region, with one popular imaging area surrounding the Northern Cross's central star, Sadr (Gamma (y) Cygni).

The star Eta (n) Cygni, sits halfway between Sadr and Albireo. Close to it is an exotic object. Cyanus X-1. a strong X-ray source and the first accepted to be directly related to a black hole. You can't see the black hole, but it is possible to see the ninth-magnitude star HDE 226868 with which it is in mutual orbit.

Surprisinaly, Cyanus only contains two Messier objects, both open clusters, M29 is located 1.8° south of Sadr and is rather underwhelming, its box shape earning it the nickname the Cooling Tower Cluster. M39 is more impressive. Sitting 2.9° north of mag. +4.0 Rho (p) Cygni, this triangular-shaped cluster stands out well against the Milky Way background. A line extended from nearby Pi-1 (π¹) through Pi-2 (π²) Cygni brings you to the Cocoon Nebula Sh2-126, associated with the embedded open cluster IC 5146. Astrophotography reveals a striking feature of the surrounding grea, the dark nebula B168 that forms a 'lane' towards the Cocoon.

Another famous patch of nebulosity sits 2.3° east of Deneb. Known as the North America Nebula, NGC 7000, this 120x100 arcminute patch of glowing hydrogen is best seen with the naked eye from a dark-sky location. It's another popular photographic target, long exposures revealing its pink-red colour and the fact that it really does resemble the outline of North America. Immediately to its west is IC5067/70, the Pelican Nebula, another popular photographic target, which resembles the profile of a pelican's head.

Switching to the western wing, you'll find the Blinking Nebula, NGC 6826. This 10th-magnitude planetary nebula gets its name because if you look directly at its central star, the star's brightness overwhelms the surrounding nebulosity, causing it to disappear, Look slightly to the side of the object using the technique of averted vision, and the nebulosity magically reappears.

▼ Two of Cygnus's







Reza Hakimi, Moomei, Iran, 10-11 July 2018

HOTSHOTS WINNER OCTOBER 2018

October

There are stunning views in October's skies as several galaxies become favourable

3 OCTOBER

Moon lies very close to Jupiter this evening

8 OCTOBER

Moon's Jewelled Handle effect visible at 17:45 UT

10 OCTOBER

Southern Taurid meteor shower reaches its peak

21 OCTOBER

Peak of the Orionid meteor shower this evening

31 OCTOBER

Jupiter sits 24 arcminutes south of the Moon's centre in daylight



The October night sky can be thrilling. In the early evening there are the constellations of autumn, led across the sky by Pegasus the Flying Horse. These are followed by the dramatic constellations of winter which dominate the morning sky

The Great Square of Pegasus is prominent early evening. This is a large asterism formed from four medium-bright stars. Despite the name, none of the square's sides have equal length. Across the top it is 12" wide while at the bottom it is 17° wide. The sides are more even at 13" on the west and 14" on the east. For comparison, your fist at arm's length is about 10" wide.

The number of stars seen within the square reveals your sky clarity. A count of zero means your sky is poor and probably light-polluted. An average UK sky should reveal at least five stars. A dozen or more indicates particularly good skies, but is way off the theoretical maximum of 35 stars. If you can count this many then your sky is quite exceptional.

The Great Square is a starting point for locating the Andromeda Galaxy, M31. Extend a line from Scheat (Beta (6) Pegasi) in the northwest corner through Alpheratz (Alpha (a) Andromedae) in the northeast.

MOON PHASES

Key stages in the monthly cycle







VISIBLE PLANETS

Where to spot the planets this month





MERCURY 20 October, but is poorly placed all month

Slowly drawing from the Sun

into the evening sky. Best seen at the end of October

Morning object slowly separating from the Sun. Telescopically rother tiny.

Observing window is closing.

Vaxina crescent Moon nearby on the evenings of 3 and

SATURN Evening planet Saturn has a close visit from the first quarter Moon

URANUS 28 October and is visible all night long.

NEPTUNE Well placed as an evening planet in Aquarius, close to the star Phi (a) Aquarii.

Keep going for the same distance again, bending north as you go to arrive at medium-bright Mirach (Beta (β) Andromedae). Turn 90° to move up the sky and reach Mu (u) Andromedae, Keep going to locate dimmer Nu (v) Andromedae and M31 is the elongated fuzzy patch just west of this star.

M31 is the furthest object you can see with the naked eve from a typical dark site, a spiral galaxy ground 2.5 million lightyears away. What you're looking at here is just the bright core of the galaxy. Binoculars or a small telescope should reveal M31's two closest satellite galaxies. Relative to the core of M31, M32 sits 0.5° to its south, while the M110 lies 1.6" to the northwest.

The hero and the blinking demon You can use M31 to locate M33 the Triangulum

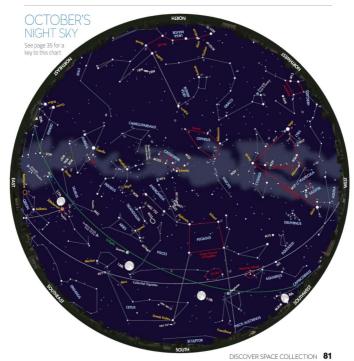
Galaxy. Extend a line from M31 through Mirach for the same distance again to arrive at M33. This is slightly further away at 3 million lightyears and much harder to see because it's virtually face on to us, resulting in low surface brightness. The core is brightest and can be seen with binoculars or a small telescope, or even with the naked eve from an exceptionally dark site.

Wedge-shaped Andromeda stretches east towards her mythological rescuer and husband, Perseus,

the Greek Hero. Its brightest member is the whiteyellow supergiant Mirphak (Alpha (a) Persei), sitting in a semicircle of fainter stars. Many of these are associated, belonging to the **Alpha Persei Cluster** or Melotte 20.

South of Mirphak lies the edipsing binary Algol (Beta (iR) Persal), also known as the Demon Star. This is a system where two mutually orbiting stars appear to eclipse one another as seen from Earth. When Algol's brighter primary is edipsed by the dimmer secondary we see a dip of 1.3 magnitudes. When the brighter star eclipses the dimmer one, the dip is so shallow that specialist equipment is needed to detect it. The primary eclipses last for around nine hours and occur every two days, 20 hours and 49 minutes.

To the north of Andromeda you'll find her mother Queen Cassiopeia, the easily identified W-shopped constellation. Drowing an imaginary line between the central star of the W. Gamma (ył Cassiopeiae, and Mirjahak in Perseus, haffway olong its length you'll find two beautiful open clusters, NGC 884 and 807, Innown as the Double Cluster. With dark conditions, binoculars or a telescope on a low power will give the best view of these stunning objects.



IC1396 >

Adam Shewan, York, August 2017 & August 2018

NOVEMBER 2018

November

A transit of Mercury and a lunar occultation of Jupiter are two rarities to relish this month

11 NOVEMBER Transit of Mercury

17/18 NOVEMBER

Peak of the Leonid meteor shower. with Moonlight interference

24 NOVEMBER

Moon and Mars appear close in the morning sky

Daylight lunar occultation of Jupiter

Venus, Moon, Jupiter and Saturn appear close in the evening sky





Novembe. at a glance

MOON PHASES

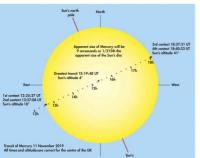
Key stages in the monthly cycle











VISIBLE PLANETS

Where to spot the planets this month



MERCLIRY Transit on 11 November, Visible from 20 November in mor

sky. Greatest western elong on 28 November

Evening planet best seen towards the end of November. Close to Jupiter on 24 November

ically poor morning planet. A thin waning crescent Moon close by on 24 November.

Evening planet joined and

utshone by Venus on 24 November, Occulted by the Mo in daylight on 28 November.

Evening planet slowly engulfed by the twilight. Moon lies near on 2 and 29 November.

URANUS Well positioned all month Located in Aries and just possible to see with the naked eve

NEPTUNE Well placed during November, able to reach highest position due south, in darkness all month.

The transition zone between the autumn and winter sky is very evident in the hours before midnight during November. The watery constellations of Pisces, Cetus and Eridanus occupy a large area of the sky, but they aren't particularly striking in shape or form.

The most recognisable part of Pisces is the asterism known as the Circlet, which lies to the west of the constellation and south of the Great Square of Pegasus. One particularly interesting thing to look out for here is the very red, fifth-magnitude 19 Piscium which marks the eastern extremity of the Circlet. Cetus is the most identifiable of the three water

constellations mentioned, thanks to the irregular pentagon shape which is supposed to represent the head of a sea monster. The brightest star in that 'head' is orange giant Menkar (Alpha (a) Ceti). The brightest star of Cetus as a whole is Deneb Kaitos (Beta (B) Ceti) and this can be found by extending the line made by the eastern side of the Great Square of Pegasus south. Located a little north and around 14.5" east of Deneb Kaitos is Tau (1) Ceti, a mag. +3.5 star 11.9 lightyears away from us, which has been the subject of many science fiction novels. This is well deserved because, innocuous though Tau appears, it's a single star much like the Sun which is known to have at least five planets in orbit around it.

A A rare transit with the low

Celestial hide and seek

The region to the north of the head of Cetus is occupied by Aries the Ram. The most prominent shape here is a bent line of stars which starts in the east with Hamal (Alpha (a) Arietis). The most interesting thing about Hamal is that, as stars go, it's pretty average. Although this may not sound particularly exciting, it does mean that Hamal is a good reference star against which to compare others. November 2019 plays host to a rare event known

as a transit of Mercury. Both of the inferior planets.



the capability of passing between Earth and the Sun. When they do so. the planets look like dark silhouettes against the Sun's disc. Transits of Venus are very rare, the next not due to occur

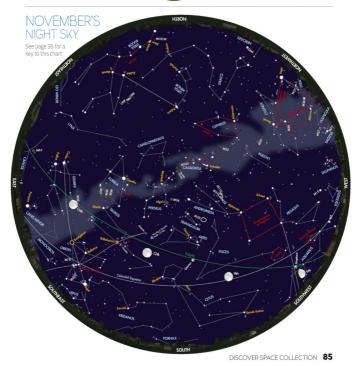
Mercury and Venus, have

■ Jupiter will be occulted 28 November, Disappearance until December 2117. Those of Mercury occur more frequently, but are still infrequent enough to attract a lot of attention. The last occurred on 9 May 2016 and, after this one on 11 November, the next one isn't due until 13 November 2032.

Lasting around five and a half hours, the transit takes place a taime when the Sun isn't particularly high as seen from the UK, and this will add an extra challenge to observing the event. In addition, from the UK it's not possible to see the whole transit as the Sun will have set by the time Mercury moves off the Sun's disc.

WARNING
Do not look directly
at the Sun with the
naked eye or any
unfiltered optical
instruments

As if the transit of Mercuny isn't enough to satiate your desire for infequent astronomical events, a rare lunar transit of a planet tokes place on 28 November. This will also be a challenging event to observe because it takes place during the hours of daylight as seen from the UK. The planet in question is Jupiter and this will be hilden behind the disc of a 3% like waxing research will be hilden behind the disc of a 5% like waxing are search was As the Moon gathers altitude it should become easier to see – although by no means child's play – giving you approximately one hour to prepare for Jupiter's emercence from behind the illuminated Junar cressent.



December

Taurus continues to delight as Orion provides ample hunting ground at the year's end

4 DECEMBER

Jupiter's disc in daylight

11 DECEMBER

Venus and Saturn close in the evening twilight

13/14 DECEMBER

Callisto's shadow transits A Moonlit Geminid meteor shower peak

22/23 DECEMBER

Favourable conditions for the peak of Ursid meteor shower

29/30 DECEMBER

Crescent Moon appears close to Venus in the evenina sky







ecembe

The magnificent constellation Orion starts to dominate the sky during December (turn to page 48 to find out more about its mysteries). The Belt of Orion is a useful signpost to other stars - follow its line southeast to arrive at brilliant Sirius, the brightest nighttime star of them all. In the other direction lies the distinctive orange Aldebaran, the brightest star in Taurus the Bull (read more about the secrets hidden in Taurus on page 90).

▼ The Ursid meteor shower



MOON PHASES

Key stages in the monthly cycle







VISIBLE PLANETS

Where to spot the planets this month





Located near Venus and Saturn

MERCLIRY Well positioned in the morning sky at the start of December. The planet remains visible until Christmas Day

Spectacular evening object setting three hours after the Sun at the end of December. Close

MARS

Mars is a morning object and NEPTUNE telescopically rather poor, Currently close to Phi (a) Aguarii showing a tiny 4 arcsecond disc.

at the start of December. Evening planet low in the southwest after sunset.

Evening planet well positioned in Aries. Best placed in early December.

Below Orion is a small pattern of stars representing Lepus the Hare, a constellation resembling the sign for infinity with ears! Lepus contains several interesting objects, including the red variable star R Leporis, known as Hind's Crimson Star. Find it by extending the line from Arneb (Alpha (a)) through Mu (u) Leporis and extending it for approximately four-fifths that distance again.

A similar star-hopping technique can be used to locate M79, an eighth-magnitude globular cluster south of the infinity pattern. Extend a line from Arneb through Nihal (Beta (6) Leporis) for the same distance again to find it, although its low altitude makes it something of a challenge from the UK.

A loyal hound and a Christmas tree The region to the west of Lepus appears rather sparsely occupied, being home to the meandering form of Eridanus the River, the longest constellation in the

night sky. From the UK we only get to see the northern part of the river: those who live further south can see the bright star Achernar (Alpha (a) Eridani) which marks the southern end of Eridanus.

There's a ghostly reflection nebula close to the star Cursa (Beta (ß) Eridani) known as IC 2118, the Witch Head Nebula. It gets its informal name because in long-exposure photographs it has the appearance of a witch's face in profile. The nebula has a low surface

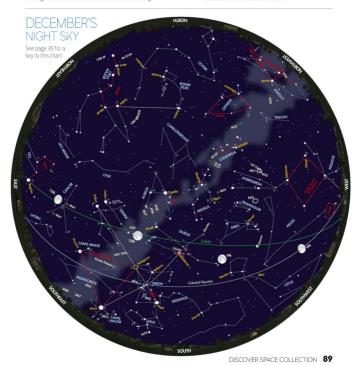
brightness and is best suited to astrophotography.

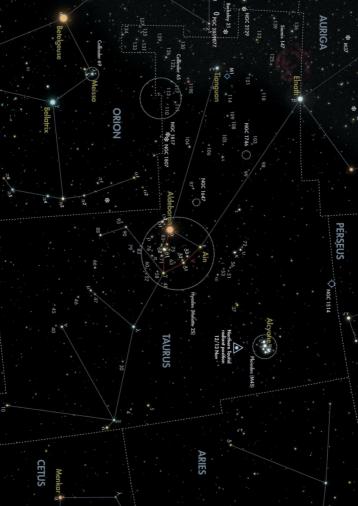
The stars to the west of Lepus belong to the lovely constellation Conia Major the Great Dog, running up the sky towards its master. Orion. The Dog's head is represented by a right-angled triangle which can be seen to the east of Sifus (Alpha (a) Conis Majoris). One front leg is marked by Mirzam (Beta (i) Conis Majoris) to the west of Sifus. The rear legs and tail are quite distinctive, although from the UK they never get particularly high above the southern horizon.

North of Canis Major is **Monoceros the Unicorn**. This is a large but ill-defined constellation full of interesting

objects. Beta (β) Monocerotis, for example, is a lovely telescopic triple star. Then there are the clusters M50 which lies 4.3" north-northeast of Canis Major's nose and M48 which lies slightly to the east of Monoceros in neighbouring Hydra. Both are binocular objects.

Further north there is NGC 2244, an open cluster which lies in the centre of the visually dim but photogenic Rosette Nebula. Slightly north of this region and on the tip of the unicom's horn is a complex region of sky which includes the Cone Nebula, Fox Fur Nebula and a rother fitting object for December, the Christmas Tree Cluster. NGC 2264.





Constellation of the season

AUTUMN: Taurus

Taurus the Bull is an iconic constellation of autumn and winter. Some suggest it may even be represented in the 17.000-year-old cave paintings in Lascaux, France, In Greek mythology, Taurus was the form Zeus took to abduct the beautiful Phoenician princess Europa.

The constellation contains two large and obvious naked-eve clusters: the V-shaped Hyades and the Pleiades, also known as the Seven Sisters. At a distance of 151 lightyears, the Hyades is the closest open cluster to the Sun. The main V-shape is large, with arms 4° in length. Orange giant Aldebaran (Alpha (a) Tauri) sits at the end of the southern arm of the 'V' but isn't part of the cluster, lying approximately half as far at 65 lightyears. The Hyades is old, around 625 million years. The star Theta (0) Tauri sits halfway along the bottom arm of the V, a naked-eye optical double.

The Pleiades are more distant at 444 lightyears. With an age of around 100 million years, this is a young cluster full of hot, bright, blue stars. The brightest star here is Alcyone (Eta (n) Tauri). The cluster contains at least 1,000 stars within its core radius of eight lightyears, and has a tidal radius of 43 lightyears. It is estimated that one quarter of its members are brown dwarfs; objects of insufficient mass for nuclear fusion to begin.

The bull's eye and beyond

Long-exposure photographs show the Pleiades full of blue reflection nebulosity, the result of the cluster stars passing through an unconnected dust cloud in the Milky Way, Under dark skies, some of this nebulosity can be alimpsed through a telescope.

The Hyades and Pleiades are large and bright enough to be part of the 'structure' of Taurus, the Pleiades marking the creature's shoulder blade, the Hyades its face. Aldebaran represents one eye, with Ain (Epsilon (e) Tauri) the other.

Extending the arms of the Hyades' V east brings you to the stars marking the Bull's horn tips. The upper and brighter one is Elnath (Beta (β) Tauri). Originally known as Gamma (v) Auriage, it's often depicted connected to the neighbouring constellation of Auriga the Charioteer.

The southern horn tip is marked by Tianguan (Zeta (¿) Tauri), a name of Chinese origin which means 'celestial gate'. Tianguan is useful for locating M1. the Crab Nebula, a supernova remnant from a star explosion witnessed from Earth in 1054 AD, an event bright enough to be seen in daylight for 23 days. A telescope is needed to see what's left of the star and it can be underwhelming, typically little more than a faint grey smudge. Astrophotography brings out more detail, revealing a delicate network of glowing filaments surrounding an interior glow. An extreme object known as the Crab Pulsar sits at the heart of the Crab Nebula. Discovered in 1968, this was the first pulsar to be linked to a supernova remnant. The explosion produced an incredibly dense neutron star. 20km in diameter, which spins and 'pulses' at 30 times a second.

Two low-rate meteor showers have their radiants in Taurus: the Southern Taurids, active from 10 September until 20 November, and the Northern Taurids, active from 20 October until 10 December, Both showers have maximum ZHRs of five meteors per hour.







BBG Sky at Night

Best of 2018

Fach month of this Yearbook starts with a winning photo from our monthly Hotshots competition. Now see some of the other superb images we received in 2018

> Moonset over the Angel D

> David Frost, Derbyshire, taken 20 October 2017



△ The Triangulum

Gary Stone, Peterborough, taken 18 November 2017



Moon montage ▷ David Ettie, Tyne and Wear, taken 10 November 2017

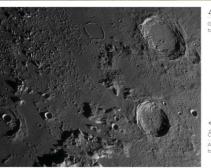


∇ The Pleiades

Jason Wiseman, Devon, taken 17 December 2017







△ Aurora

Gill Williams, Ylläs, Finland, taken 19 January 2018

& Eudoxus

Avani Soares, Canoas, Brazil, taken 25 December 2017



→ Slaley skyscape

Gary Pescod, Slaley Forest, Northumberland National Park, taken on 13 February 2018





√ The Horsehead Nebula

Simon Todd, Haywards Heath, taken 9, 11 & 15 February 2018

David Pickles. Northants, taken 14 April 2018

 ∇ Antares region

The Carina Nebula D Rafael Compassi, Presidente Lucena, Brazil, taken 23 April 2018 Mario Richter, Namibia, taken 12 May 2018



The Milky Way D

Mark Pelleymounter, Kimmeridge Bay, Dorset, taken 20 May 2018



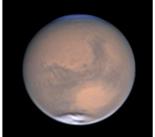




Nebula

Luis Fernando Parmegiani, Padre Bernardo, Brazil, taken 12 July 2018





 Δ Mars

Agapios Elia, Nicosia, Cyprus, taken 4 August 2018

Chris Pomeroy, Cyprus, taken 3 August 2018



Project 1: Spot the planets

Turn sky detective and see if you can pick out the unique defining characteristics of every planet in our Solar System

The planets are a constant source of wonderment, each with its own defining qualities. This project highlights some of these characteristics and challenges you to identify two aspects of each planet's appearance. One will be a moderately easy characteristic and the other harder. To help locate your targets, you can find summaries of the planets' positions in the month by month guide in this Yearbook

We'll start with inner planet Mercury, the smallest of the main planets, with a diameter of 4,879km. Its average distance from the Sun is 57.9 million km and it takes 88 days to complete each orbit. It's not the easiest planet to spot, oscillating between morning and evening sky but never straying too far from the Sun.

Our easy characteristic for Mercury is just to spot the planet in either the evening or morning twilight. Trivial as this seems, it's surprising how many amateur astronomers have never seen it! During 2019, it's best seen in the evening sky late in February into early March, and in the morning sky in mid-August and late November into early December

For Mercury's hard characteristic, once you've located it we want you to look at it with a telescope. Mercury is an inferior planet, meaning its orbit is smaller than Earth's. The angle between Mercury,

WARNING Take great care if attempting to view Mercury through a telescope if the Sun is above the horizon

the Sun and Earth means it can show a phase. Spot this and you'll have observed the hard characteristic On 11 November 2019, Mercury giving you the opportunity to catch it at a rare 'new' phase.

transits the Sun (see page 84),

the eveniece, its bright atmosphere often appearing uniform in appearance. But this isn't true and once your observing eye kicks in you should start seeing subtle shading differences. For our hard characteristic we want you to observe and record these shadings.

The next observable planet is Mars. This can be frustrating because it only presents a decent-sized disc a few months either side of opposition. This is the time when surface detail is easiest to see. The next opposition occurs on 13 October 2020. For our easy characteristic we want you to look at Mars through a telescope. When away from opposition - as is the case throughout 2019 - its disc will appear small but still shows a



◆▲ Observing the use of a portable telescope setup as the planet is often





Like Mercury, Venus is also an inferior planet, appearing to flit back and forth either side of the Sun. However, its larger orbit means its apparent separation can place it in much darker skies further out on the edge of twilight. Its thick dense atmosphere reflects sunlight well and it's the brightest of the main planets visible from Earth. Like Mercury, Venus shows phases. For the easy characteristic we simply want you to observe Venus through a telescope and note its phase.

The Venusian alobe can look uninteresting through



- ▲ The changing phase of Venus
- ◄ Planet Mars will not be at its best in 2019, appearing quite small through the eveniece of a telescope





▲ Saturn's low altitude makes spotting the narrow Cassini Division that bit harder







▲ Neptune is the furthest and dimmest of

characteristic salmon-pink colour. Despite having a larger (superior) orbit to Earth. Mars can show phases, For a bonus observation, see if you can spot this phase.

The hard characteristic is easy near opposition, but will be tricky during 2019. We want you to try to see surface detail. Mars has dark features - exposed rock - and light deserts. It also has bright polar caps. If you can spot

■ Jupiter, shown here with south up as it would be seen through the eyepiece of an inverting telescope, shows a wealth of detail, including the Great Red Spot

anythina like this durina 2019 you're doing well!

Jupiter is a gas giant with a complex atmosphere segregated into bands and zones by the planet's fast rotation. Our easy characteristic requires you to spot several features related to the planet's equator. Its fastrotating gaseous globe appears flattened into an oval bulaing at the equator, Two wide, dark bands encircle the globe, running parallel to the equator. These are the north equatorial belt (NEB) and south equatorial belt (SEB). Finally, take note of the four largest 'Galilean' moons of Jupiter, lo. Europa, Ganymede and Callisto orbit close to the planet's equatorial plane and appear as bright dots more or less in line with the equator.

For Jupiter's hard characteristic we want you to try to observe the persistent storm known as the Great Red Spot (GRS). This is always present, nestled into the southern edge of the SEB. However, its visibility isn't augranteed because you'll need to catch the planet at a time when the GRS is on the side facing Earth.

Track the ice aiants

Saturn's easy characteristic is satisfyingly predictable. We want you to observe and note the ring system. The hard challenge also concerns the rings. For this we want you to try to spot the Cassini Division, a narrow gap running between the outer A-ring and the inner B-Ring. It's only 4.700km wide and requires good atmospheric stability to be able to record it all the way around.

The ice giants Uranus and Neptune are trickier in as much as they are so distant that they don't offer much in the way of detail that amateurs can see. In both cases, the easy characteristic is to note each planet has a disc and to record its colour. For Uranus's hard characteristic we want you to see whether you can detect any banding or shading on the planet's disc. Unlike the other hard characteristics, this one is particularly tough and will require a large telescope.

A large aperture will help with Neptune too, because here we want you to try to spot its largest moon. Triton. Although Neptune is a very distant world. Triton does stand out surprisingly well through larger scopes.

How did you do?

How difficult did you find locating all the planets? How did you fare with observing their harder characteristics?

Project 2: Cracks in the Moon

Best seen under low light conditions, these lines in the Moon's surface are a fascinating feature. Use our guide to see how many you can find

The Moon is a captivating target for amateur astronomers and even a small telescope can reveal its craters, mountains, valleys and lava-filled seas. But the Moon's surface also has a lot of cracks. Many are just a few kilometres wide but tens or hundreds of kilometres long. Some are isolated, while others appear connected in delicate criss-crossina networks. Oblique lightina can make these cracks appear very dramatic and it's an intriguing exercise to trace their paths across the Moon's surface.

A lungr crack is known as a rille, from the German word for 'groove'. The Latinised form is rima for a single rille or rimae for multiple rilles. There are three different types of rilles: straight, sinuous and arcuate.

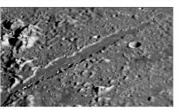
Straight rilles often look like roads from above. A graben is an example of a straight rille, a region of lunar surface which has dropped between two parallel fault lines. A classic example of a graben is the Vallis Alpes. Although too large to be considered a rille in its own right, Vallis Alpes does have a narrow crack that runs

down the middle of it. With a width of around 0.7km. this is quite challenging to see and requires really steady conditions and at least 250mm of aperture. It's a little easier with high-resolution imaging, but still far from a pushover.

Another excellent example of a straight rille can be seen running from the western rim of Hesiodus through into Palus Epidemiarum in a northeast to southwest direction. Called Rima Hesiodus, this really does look like a road crossing the lunar lava. It measures 3km in width and is around 300km in length.

Straight, sinuous and arcuate

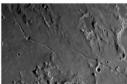
The region south of Mare Serenitatis also has some excellent examples of straight rilles. Rima Ariadaeus runs between Mare Vaporum and Mare Tranquillitatis. This is shorter than Rima Hesiodus at 220km long. but wider at 7km. It also has the classic road-like appearance of a graben. It ends close to another famous rille, known as Rima Hyginus. This is again

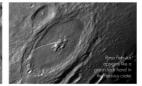


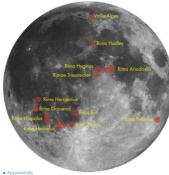
◆ The narrow rille running along the base of Vallis Alpes will require steady conditions to discern

> ◄ Rima Hadley is an excellent









lunar rilles we

220km in length, but narrower with a width of 4km. It's also rather odd, apparently changing direction when it encounters the flat floored 10km crater Hyginus which lies at the centre of the rille's run. Large apertures or high-resolution imaging show Rima Hyginus to be guite unlike a graben, but rather formed from a multitude of tiny craterlets along its length. One theory is that this was a lava tube which collapsed to form the rille, the craterlets appearing at the same time as the collapse.

A relatively short hop of 140km southwest from Hyginus brings you to 27km Triesnecker, A complex and delicate set of interconnecting cracks appear to run roughly north south to the east of this crater. The Rimae Triesnecker complex is a alorious sight when the sun is low in the lunar sky. Their width is around 2km.

The early waxing crescent or early waning aibbous are ideal phases for seeing another famous straight rille known as Rima Petavius. This 2km-wide, 80km-long crack runs from the central mountain complex of the 177km crater Petavius towards the southwest rim. This makes it look like a giant clock hand. Larger instruments should be able to make out a continuation of the crack running north from the mountain complex toward the northern rim

Sinuous rilles form when lava runs beneath the lunar surface, creating a structure called a lava tube. As time goes by and everything cools, the empty tube may collapse, revealing a meandering river-like rille. The most famous example requires at least 200mm of aperture to spot, along with good seeing. Known as Rima Hadley, this lies in a small area of lava to the east of Palus Putredinus which itself is located to the southeast



- ◄ Picking out Rima good seeing
- ▼ Gassendi crater rilles, and Rimae



of 83km Archimedes. Rima Hadley is 2.5km wide and meanders in smooth curves for approximately 80km. The 6km crater Hadley C sits more or less at its centre. The Apollo 15 mission landed close to a prominent northern bend in the rille and sent back images which showed Rima Hadley to be a gently curving depression approximately 300m deep.

Another difficult to observe set of sinuous rilles is located north-northeast of 111km Gassendi, itself a crater filled with straight rilles known as Rimae Gassendi. Rimae Herigonius are a set of 1km-wide sinuous rilles which run to the west of 15km Herigonius, towards the eastern rim of lava-filled 119km Letronne. A smaller example can be found to the west of the famous Straight Wall fault, Rupes Recta, Rima Birt is a 1.5kmwide sinuous rill that runs for 50km, A 300mm scope is required to see it convincingly.

The final type of rille is an arcuate rille. These are curving rilles most commonly found running parallel to the edge of large circular seas. One of the best examples of arcuate rilles is Rimae Hippalus. These are wide rilles which appear to run parallel to one another concentrically with the Humorum basin. A smaller example of arcuste rilles can be found inside the rim of the 98km Pitatus which is located on the southern shore of Mare Nubium, A large aperture will show a set of 1km rilles running concentrically ground the edge of the crater's floor.

How did you do?

How many of these cracks did you manage to identify and did you successfully image any?

Project 3: Seeing red

Test your star-hopping skills and see whether you can track down our selection of seasonal red beauties



In modern times, we can locate thousands of celestial objects at the touch of a button. But there's nothing quite like finding stuff manually. For this challenge, we're using star hopping to find some understated objects of astronomic beauty: red stars.

Many stars are described as red. The red supergiant Betelgeuse (Alpha (a) Orionis) is a good example, but may disappoint when you realise that it doesn't actually look red. To the uninitiated. Betelgeuse seems to have hardly any colour at all. It's only when you move your gaze between Betelgeuse and Rigel (Beta (B) Orionis) that you realise that the colour of Betelgeuse does indeed lean toward the red end of the spectrum. However, rather than being pure red, Betelgeuse has an orange hue. This begs the question: are there any truly red stars? The answer is an unsatisfying 'nearly'. A star's colour reflects its photospheric or 'surface' temperature. Higher temperatures shift a star's peak output toward the blue end of the spectrum. These stars look white. Cooler stars have emission peaks in the vellow-red part

▲ Photographing areat way to bring T Lyrae's trail on



▲ You can locate alluring red giant R Leporis near the 'ears' of the



▲ With a brightness varying over the course of a year, the strong red hue



▲ For one of your summer targets, try to find V Aquilae, a carbon star whose vivid red hue should be obvious even with binoculars



A It will require a telescope, but locating bright, pulsating Y Tauri off the tip

of the spectrum and our eyes cope better with these, seeing them as orange. One way to quantify a star's redness

is to calculate its B-V index, derived by determining the star's apparent magnitude through a blue (B) filter and a visual (V) filter. Red stars have a positive B-V index: the more positive the value, the redder the star. Our Sun's B-V index is +0.66. while that of Betelaeuse is +1.85.

The variable R Leporis, or Hind's Crimson Star, varies between magnitudes +5.5 to +11.7 over a period of 427 days. Its redness is most intense when dimmest, its B-V reaching an impressive +5.7. Through the eyepiece, R Leporis has a truly beautiful deep orange hue. It was this that caused the astronomer John Hind to describe the star as looking "like a drop of blood on a black field". The B-V decreases to +3.5 when R Leporis is brightest, but this too is quite orange.

magnitudes +7.5 and +9.2 and has an impressive peak B-V of +4.38

The spring brings new offerings and our pick is SS Virginis. This is located in the Bowl of Virgo. Imagine the line between Zaniah (Eta (n) Virginis) and Vindemiatrix (Epsilon (e) Virginis). SS Virginis is located one-eighth the way along this line, starting at Zaniah. This is another semi-regular variable, swinging between magnitudes +6.0 and +9.6. Its B-V is +3.86.

Next up. Y Canum Venaticorum is easy to find, Draw a line from Cor Caroli (Alpha (α) Canum Venaticorum) and Chara (Beta (β) Canum Venaticorum). Turning 90° to head north-northeast towards the Plough for the same distance again brings you right to it. This is a bright and deep orange beauty. Its magnitude varies from +4.9 to +7.3, with an average B-V of +2.54. Y Canum Venaticorum is also known as La Superba, a name bestowed by 19th-century astronomer Angelo Secchi in appreciation of the star's beautiful colour.

Our first summer selection is mag. +7.6 T Lyrge, with an impressively high B-V of +5.10. Locate it midway between the double star Zeta (E) Lyrae and Kappa (K) Lyrae. Next is the semi-regular variable V Aquilae, which varies from +6.6 to +8.4. This has a B-V of +4.00 and is easy to locate. Simply extend the line from Eta (1) Scuti through 12 Aquilae for 60 per cent as far again.

Finally, our autumn spots, UU Aurigae is a semiregular variable with a peak magnitude of +5.5 and a B-V of +3.10. Find it by extending the line from Hassaleh (lota (i) Aurigae) through Theta (fi) Aurigae for half as much again. Y Tauri has a magnitude range of +6.5 to +9.2 and a B-V of +3.44. This is another easy find, located midway between Zeta (¿) Tauri and Chi-1 (¿¹) Orionis.

On red alert

To get you seeing red and honing those starhopping skills, we've picked a couple of high B-V stars from each season, starting

with R Leporis. This is easily found by extending a line from Arneb (Alpha (a) Leporis) through Mu (u) Leporis for 80% that distance again - an easy star hop so long as you can recognise Lepus! Our second winter choice is the semiregular variable RY Monocerotis. Draw a

line from Sirius (Alpha (a) Canis Majoris) through Theta (t) Canis Majoris, extending the line for the same distance again. RY Monocerotis sits 1° east-southeast of this position. Use a low-power eyepiece and the star should be easy to find, thanks to its colour, RY Monocerotis exhibits variation in brightness between

A "Like a drop of a telescope

▲ Stunning, deep

How did you do?

Did you track down all of the red targets we featured? Which did you think was the reddest?

How to

Collimate your binoculars

Bring your binoculars back to life with Stephen Tonkin's step-by-step guide to correcting misalignment

The increasing availability of inexpensive astronomical binoculars has opened this side of our hobby to tens of thousands of people who would otherwise have been priced out of it. But with that have come some compromises, one of which is that low-cost binoculars can easily lose collimation.

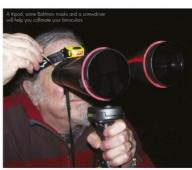
Collimation, as far as binoculars are concerned. means that the images from the two optical tubes must merge within very tight tolerances. The night sky is very demanding of optical systems, so a slight misalignment that you may not notice in daylight can become especially apparent under the stars. If your binoculars are badly out of collimation they'll give a double image which is, at best, very annoying. However, even if collimation is only slightly outside acceptable tolerances, your brain will attempt to compensate for it, putting strain on the muscles that move your eyes, which can lead to headaches or nausea

If your binoculars arrive out of collimation or if they're still under guarantee, you should return them to the vendor to be remedied, because the measures described here will immediately invalidate any warranty. If the guarantee has expired, you may wish to have them collimated by a professional repairer. but for many inexpensive binoculars this will cost more than the price of replacing them. In this case, you have nothing to lose by attempting the job yourself.

It takes specialist equipment and a lot of skill to perform a full collimation, so what we'll do here is what's known as a 'conditional alignment', where we align the optical tubes with each other in order to get a single image, but we won't attempt to align them with the hinge as well. As a result, they'll only be in good alignment for the interpupillary distance at which you align them; this is the 'condition' in conditional alignment.

Aligning the prisms

The typical cause of Porro-prism binoculars losing collimation is being dropped or receiving a knock that shifts one of the prisms. In the most common form of inexpensive binoculars, each prism is held in place with a spring clip that tensions it against a screw that tilts the prism. This arrangement means the prisms are vulnerable to being dislodged by impacts, but it also



means that you can usually correct the problem by adjusting the tilting screws.

First of all, thoroughly examine your binoculars to make sure that any miscollimation isn't due to external physical damage, such as an objective tube being bent out of alignment. Fixing this kind of defect is beyond

Tools and materials



Print out the Bahtinov masks on transparency film suitable for your printer. This removes the need to cut out the individual slots. Leave three tabs evenly spaced around each mask to tie down on to your binoculars Anaglyph glasses are a useful alternative to Bahtinov masks and can be used in daylight.

A head-torch is essential for hands-free illumination of the task. An L-bracket or hinge clamp is essential to hold your binoculars still while you collimate them. Select a small flat-head screwdriver to fit the collimation screws. Use elastic bands or sticky tape to secure the Bahtinov masks to your binoculars.

the scope of this guide. Next, you should test the binoculars under a night sky. To do this, you have to overcome the ability of your eyes and brain to merge images that aren't actually merged.

There are three ways to fool your brain into thinking that the separate images are of different objects. each one offering an increasing level of complexity and precision.

The first and most simple method is to defocus one side of your binoculars. The focused image from the other side should be central in the defocused image.

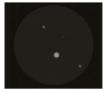
The second method is to use different-coloured filters, such as 3D anaalyph alasses (although coloured cellophane sweet wrappers will do in an emergency),

so that the image from each side is a contrasting colour. You can also use these during daylight to merge the images of, for example, a distant TV antenna.

The third and most precise method is to use Bahtinov masks, orientated at right angles to each other. The simplest way to obtain these is to use a mask generator, such as astrojargon,net/ maskgenerator.aspx. Note: the website will request the focal length of your objective lens. If you don't know this, assume it's four times the aperture.

If you follow the step-by-step guide below and get the collimation as precise as you can, you'll give your binoculars a new lease of life and be able to enjoy many hours of strain-free observing.

Step by step



Step 1 The first step is to determine whether your

binoculars are really out of collimation. Observe a bright star and defocus the right-hand eyepiece. Any displacement of the focused star from the centre of the defocused one is the error you need to correct.



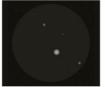
Step 4

The collimation screw tilts the prism against the tension of the spring clip. Look through both eyepieces, and rotate each screw by no more than one eighth of a turn at a time to see what effect it has on the image, then return the screw to its original position



Step 2

It's a good idea to locate the collimation screws in good lighting. They will be small and may be covered by casing and adhesive. Remove material that's covering them. Then choose the correct size of flat-head screwdriver to turn the screw and break any locking adhesive.



Step 5

Make sure that you have set your binoculars to your interpupillary distance (IPD) then, using what you've learned about the action of each collimation screw, merge the images into one, again using no more than one eighth of a turn of the screw each time



Step 3

Even if you don't normally mount your binoculars, you should do so now. Polaris, if you can see it, is a good target for collimation as it doesn't move appreciably. If you choose a terrestrial object, pick one that's at least a kilometre distant.



Step 6

When you think you have the images merged, check the IPD again and use the Bahtinov masks or anaglyph glasses to make sure that they're as closely merged as you can get them. Lastly, use them normally and enjoy the new and improved view!

Build a low-cost wooden tripod

Make a sturdy, portable alternative to a solid pier or home observatory with Mark Parrish's DIY guide



Most astronomers would love a permanent observing setup in their garden, and to that end an observatory or a solid pier is often the dream solution. Of course, that's not always possible. What if there are surrounding trees or buildings that mean you need to shift your setup to different parts of your garden to view different parts of the sky? What if you can't make permanent changes to your property?

Whatever limitations you face, the inevitable result is that you are obliged to use a moveable tripod. This brings its own issues. Taking your tripod indoors after each session loses your carefully gained polar alignment, but leaving what can be a very expensive piece of equipment outside could be a big attraction for thieves. That's why we're going to explain how to build an inexpensive, stable tripod that can be left in place or moved around as needed.

The timber you'll need can be readily acquired from a local builder's merchant. We chose 'prepared' (planed) timber because it is nicer to handle and looks attractive. but you could use equivalent rough-sawn wood to

reduce the cost further. The bolted construction of the lea frames and the use of hinges to join the frames mean you don't need to have any special skills to assemble the main parts. The tripod is immensely strong and has plenty of mass, which helps to provide a very solid platform on which to fix your mount. Our tripod is 1,065mm (3.5ft) tall, but you can alter the design of the central pier; for example, make it taller if you have a long refractor or shorter for a big Newtonian.

We used a dark brown preservative wood stain, but there are many colours available, so you could camouflage your tripod to blend in with the bushes or match other garden furniture and features.

The top of the tripod and the mount fixing consist of a number of hexagonal plywood plates. Luckily, if you can't remember your school geometry lessons, there are downloadable templates available at https://bit. lv/2AQv5fF. It's best to cut them simultaneously so that each ends up the same shape. If you mark one edge of the stack you can realign them easily even if they gren't perfectly symmetrical.

Platina up

Your telescope mount will be bolted to its own removable hexagonal top plate, which fits precisely into the tray-like, fixed top plate on the tripod. This

Tools and materials



Hand saw, jigsaw or chop saw, drill and hits

MATERIALS

Two 8ft lengths of 98x38mm construction timber. Four 8ft lengths of 63x38mm construction timber. 1/8th of a sheet of 18mm plywood or similar, plus a few small offcuts of plywood.

SUNDRIES

Nine M10x150mm coach bolts with nuts and washers. Three pairs of 75mm butt hinges and screws. A few each of 4-inch and 2-inch woodscrews. A pair of M6x60 bolts with wingnuts.

Approximately 0.5l of preservative wood stain.

means you can lift off the whole assembly after an observing session and then replace it the following evening, maintaining your polar alignment. We drilled two holes through our top plates and used a couple of small bolts with wingnuts to hold them together once the mount assembly is lifted on, so there is no danger of equipment being knocked off during use. You could also use a G clamp for this purpose.

Adapting the removable top plate to suit your particular mount may require a bit of careful measuring and cutting but, having examined a multitude of EQ mounts from various manufacturers, we concluded that you will normally just have to cut a large round hole (around 60mm in diameter) in the upper, thick

layer of plywood and a smaller 10mm hole in a thinner layer that is subsequently glued beneath, Many manufacturers sell 'pier top adaptors' for just this kind of project, and they are worth investigating if you prefer to keep things simple.

If you need a stable, non-permanent, economical base on which to mount your equipment, we hope that you will want to give this project a try. You may have a number of mounts, in which case you could make a separate top plate for each one, so swapping them becomes a cinch. You may even decide to make a couple of tripods, locating them in the optimum greas of your garden for different sections of the sky and moving about during a night's observing.

Step by step



Step 1

After carefully marking out the parts, use a saw to cut them to size. You can use the downloadable templates at https://bit ly/2AQy5fF to help you get the angles right. If you have access to a chop saw you could save some time here. Check the similar pieces match.



Step 4

Bolt the three triangular leg frames together. Each short 'foot' section should be fixed between the two leg struts with long woodscrews once you have aligned the other elements. Fix a pair of hinges to each frame, then join all three frames to form the tripod.



Step 2

Drill the 10mm holes in one example of each part. Use this to transfer markings to the corresponding parts. You can line up the parts with the 'master' on top, then put a bolt in the hole and strike it with a hammer, denting the one below



Step 5

Mark out the plywood hexagons. Drill a centre hole in each and a second hole near to one corner, then use small bolts or dowels to hold the layers together before cutting - this ensures replica shapes. Screw one fixed top plate to the tripod and add plywood lips



Step 3

Once all the timber is cut, sand any rough edges and check it all fits together. If the holes are tight you can 'waggle' the drill or re-drill them using a 10.5mm bit. Now is a good time to paint the wood with preservative stain



Step 6

Measure your mount. Cut a suitable hole in a removable top plate. Glue a block of hardwood or similar into a notch to form a pin for the alignment screws. Screw a second thin layer of ply (with a central hole for the fixing screw) to the underside.

Build a rolling scope platform

You get your setup just how you like it... and then it's time to pack it away again. Mark Parrish has the solution, with this simple rolling platform



to handle and can be painted to match your other equipment. The plywood parts provide bracina: we made ours from 9mm thick sheet, but you could use a slightly different thickness if you already have some available. The dolly frame needs to be rigid, but as most of the scope's weight is supported close to the wheels it doesn't need to be too heavy. Individual parts are bolted together, allowing you to take the dolly apart for easy storage or transport, Furthermore, the design allows for resizing should you ever decide to upgrade your tripod.

Start by looking at the downloadable plans available at https://bit.lv/2DpJIC7. They show dimensions for our example dolly, and these measurements should suit most small to medium tripods. We have indicated the positions we drilled holes in the bracing plates, but we'd recommend lining up your beams and tripod support plates with your own tripod to determine the best position for these holes before drilling. Ideally, the wheels and jacking points should be as close to the tripod feet as possible.

For wheels, we chose 150mm plastic ones with rubber tyres. These run well on M12 studding axles and they seem to cope with regular gravel drives and garden lawns. Because the speeds involved are low and distances relatively short, pneumatic tyres and

Most medium-sized mounts and telescopes are too heavy and unwieldy to move once they've been assembled, which means setting up and packing away each night can become a slog. There's the numerous journeys back and forth to the garden, for one. Refitting mount heads and weights takes up time when you could be out observing, and rebalancing a telescope can be fiddly work in the dark. It is also all too easy to misplace leads and small parts. This project, an inexpensive rolling dolly, provides

a workaround. Leave your fully assembled setup on top of it and whenever the skies are clear you can simply roll the whole unit into position. If the weather turns against you, you can be back in the warm in next to no time, secure in the knowledge that nothing is left behind

The timber for the dolly's axles can be readily obtained from a builder's merchant. We chose 'prepared' (planed) 45x45mm timber because it is nice

Tools and materials

Handsaw or jigsaw, hacksaw, drill and bits (for 12mm and 8mm bolts). spanners, screwdriver.

MATERIALS

Approximately 3m of 45x45mm construction timber, 1,200x600mm sheet of 9mm plywood, length of suitable timber or rope for the

Four 150mm-diameter wheels with rubber tyres, 1m of M12 studding (diameter to suit wheels) with nuts and washers, six M8x75 bolts, 75mm butt hinge, epoxy resin glue, 25mm woodscrews,

Preservative wood stain or paint.

How to...build a rolling platform

proper bearings are not necessary unless you are scaling up the design for a heavyweight mount or you have particularly bumpy around to cover.

Finessing the design

If your tripod legs have slightly pointed feet, a simple hole drilled in the lea support plate will help locate them. If you have a tubular design, you could make L-shaped corner blocks and screw them into each plate to prevent your tripod from sliding off. Tying your tripod down with straps or bungees is also a good idea. in case you bump against something or have to stop suddenly. The last thing you want is for your setup to pitch forward and crash to the ground.

On the front of the dolly is a rigid towing handle. It attaches to the front of the steering mechanism via a hinge, so it can be folded back against the tripod when not in use. You could also use a loop of rope or webbing as a simpler alternative. When you have your dolly in position, screw down the jacks to stop it moving accidentally. Once you are polar aligned, you could also mark the jack positions on the ground to aid repositioning on future nights.

We also found that the bracing plate below the tripod is ideal for holding accessory cases and a power tank, but you could consider customising this area to provide racks and boxes for your accessories and imaging equipment.

Step by step



After carefully marking out the parts, use

a saw to cut them to size. You can tape multiple

plywood parts together and cut in one go so

that they match each other. If you have access

to a band saw or jigsaw you could save some

Step 2

With your main parts cut, drill holes in the plywood parts first, then line them up with the axles. Place your tripod on top to check size. Mark through the holes to find out where to drill into the axles. Do this, then assemble the timber and plywood parts.

Step 3

Make the steering 'box'. It's held together with 25mm woodscrews inserted through the plywood plates into the axle and front blocks. A piece of studding provides a pivot. Use Nylock nuts and washers and make sure it turns smoothly but is not wobbly.



Step 4

Step 1

time here

Cut lengths of studding for the axles using a hacksaw. Allow enough for a nut and washer either side of the wheel, plus about 50mm to be glued inside the wood. File the ends of the studding so that they aren't sharp and the nuts fit on nicely.



Step 5

Check that the wheels spin freely without wobbling. Use epoxy resin to glue each axle in, taking care not to get any glue on the moving parts. If the wheel needs to be removed later on, you should be able to simply undo the outer nut



Step 6

Once the tripod foot plates and jack blocks are screwed into position, drill through both for the studding. An enlarged hole on the underside gives you room to glue in a captive nut (normal type). Use a Nylock nut (or a wooden disc) to make the upper turning knob.

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Sky at Night

Glossary

Essential astronomical terms explained

The diameter of a telescope's light-gathering main lens or mirror. The bigger the aperture, the brighter and more detailed the image.

AVERTED VISION

A technique for observing faint objects through a telescope. It involves viewing slightly to the side of the object, allowing its light to fall on an area of the eye more sensitive to light.

BINARY STAR

A system of two stars that orbit a common centre of gravity. The brighter star is classified as the primary, while the dimmer is the secondary.

BLACK HOLE

The collapsed core of a massive star. Its concentration of mass is so dense that nothing, even light, can escape its gravitational pull.

CELESTIAL EQUATOR

A projection of the Earth's equator into space.

CELESTIAL SPHERE

The projection of the night sky on to an imaginary sphere around the Earth. The astronomical coordinates of right ascension and declination are also mapped onto this sphere.

COLLIMATION

The process of alianing the optical elements of a telescope or binoculars

COMET

A ball of frozen gases, rock and dust that orbits the Sun

DECLINATION

The celestial equivalent of latitude, this is the distance of a body north or south of the celestial equator, measured in degrees, minutes and seconds. Objects north of the celestial equator have a positive declination, while those south of the celestial equator have a negative declination.

The apparent annual path of the Sun against the background stars.

ELONGATION

The angle in the sky between a planet and the

- at its farthest apparent distance from the Sun. The two times each year - on or near March 21 and September 22 - at which the Sun crosses the

EQUINOX

celestial equator

Sun, as seen from Earth. At its greatest elongation

(east or west), a planet is - from our perspective

FOCAL LENGTH

The distance between a telescope's main lens or mirror and the point at which an image is brought into focus. You can work out magnification with this number; this is calculated by dividing the focal length of a telescope by that of an eyepiece.

GIRROUS

When the Moon or other body appears more than half, but less than fully, illuminated (from Latin gibbus, meaning 'hump').

LIBRATION

The apparent tilting or wobble of the Moon as it orbits Earth, allowing 59 per cent of its surface to be seen over a period of time.

LIGHTYEAR

A unit of measurement equal to the distance that light travels in one year - about 5.8 trillion miles.

MAGNITUDE

The brightness of an astronomical body. The lower the number, the brighter the object. Magnitudes brighter than zero are represented with a negative number.

MERIDIAN

An imaginary line circling the Earth from north to south that marks the point at which the Sun is at its highest in the sky.

METEOR

Also called a shooting star, this is a small particle of dust or rock that burns away on entering Earth's atmosphere. A meteor shower is when meteors are observed to radiate at regular intervals and from a particular region in the sky as Earth passes through debris left by a disintegrating comet.

NEBULA

A cloud of interstellar dust and gas, usually lit up by stars inside or nearby.

OCCULTATION

When the Moon or a planet conceals or obscures a more distant planet or star.

OPPOSITION

The position of a superior planet when it's opposite the Sun in the sky.

PHASE

The apparent change in shape of a celestial body as different amounts of its surface are illuminated by the Sun. The Moon displays four main phases: new, first quarter, full and last quarter.

RIGHT ASCENSION

The celestial equivalent of longitude, right ascension is measured in hours, minutes and seconds. The Oh line is measured from the point where the Sun crosses the celestial equator on its apparent path through the sky each year, a point called the vernal equinox.

STAR

A giant ball of hot gas that creates radiation through nuclear fusion. Stars are classified by their spectra (the elements that they absorb) and their temperature. The seven main types, from the hottest to the coolest, are: O. B. A. F. G. K and M.

STAR CLUSTER

A grouping of stars, from a few dozen to a few hundred thousand, bound together by their mutual aravitational attraction.

TRANSIT

The passage of a celestial body across the disk of a larger one, such as Mercury crossing the disk of the Sun or Ganymede's shadow moving across the surface of Jupiter.

UNIVERSAL TIME (UT)

A timescale based on the rotation of the Earth on its axis, UT is measured from midnight at the Greenwich Meridian, In astronomical use, UT replaced Greenwich Mean Time (GMT), which was measured from the Greenwich Meridian at midday on 1 January 1925.

VARIABLE STAR

A star that fluctuates in brightness over time, such as eclipsing binaries like Algol or pulsating stars like the Cepheids.

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